



Spectropolarimetric study of the cool RV Tauri star R Scuti

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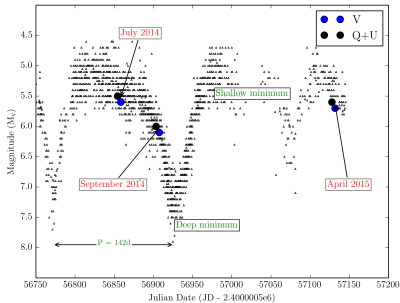
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Abstract

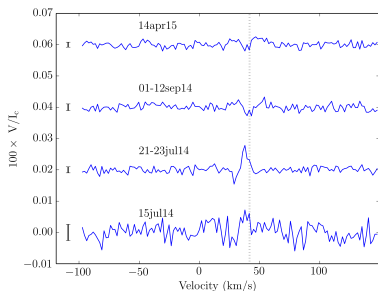
Context: With NARVAL@TBL we have initiated in spring 2015 a 2-years campaign dedicated to a sample of cool and evolved stars including pulsating RV Tauri stars. We intend to spectropolarimetrically monitor R Scuti, the prototype of such variable post-AGB targets along its pulsation period.

Aims: We want to confirm surface magnetic field detection with the circular polarization (V) and study the linear polarization (Q/U).

Results: In the first part we confirm the detection of a magnetic field about $(0.9 \pm 0.5 \text{ G})$ at the luminosity maximum of the star. In the second part, we show our first investigations about linear polarization.



Observations of R Sct along its pulsation period.



Evolution of circular polarization. The gray dotted line depicts the Radial

Velocity (RV) of the spectral line.

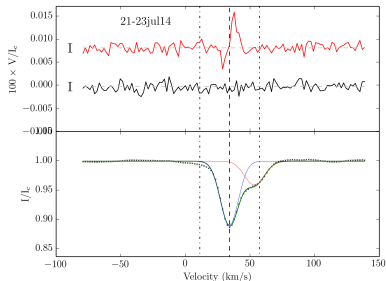
■ Multi-line extraction (LSD) on multiple spectropolarimetric sequences:

- S/N = 40 000 – 120 000
- Definite Detection at luminosity maximum (21-23 July 14)

■ Longitudinal field measurement:

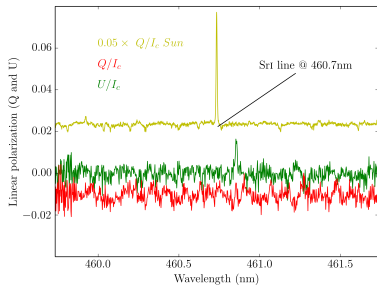
$$B_{\ell} \propto \frac{\int vV(v) dv}{\int [1 - I(v)] dv}$$

- With full I line: $B_{\ell} = 0.9 \pm 1.8$ G
- Blue component only: $B_{\ell} = 0.9 \pm 0.5$ G



Dashed line: Central RV of blue component.

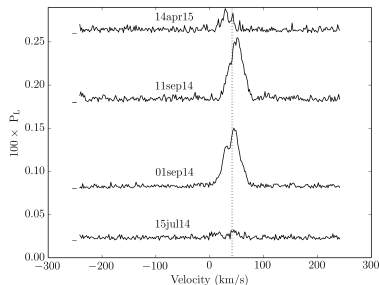
Dot-dashed lines: Integration bounds for B_{ℓ} .



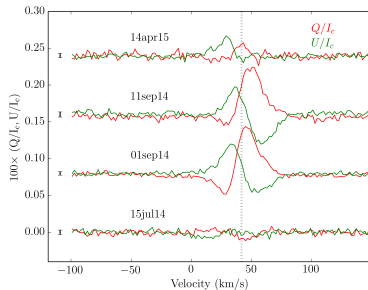
The polarization of the Sr I line in the 2nd solar spectrum (Stenflo 1914) and R sct.

- Strong linear polarization in some $\uparrow$ individual spectral lines
- Average (of about 1000 lines) signal $\uparrow$ detected with LSD
- ➡ Strongest after luminosity maximum
- The shape of Q and U remains constant $\rightarrow$

- ➔ Systematic phase-polarization relation?
- ➔ Physical origin of linear polarization?
- ➡ Unrelated to magnetic field?



Evolution of linear polarization with the phase.



Evolution of linear polarization LSD profiles, with the phase.