## Solar wind dynamics in the environment of comet 67P

IRF



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**CONTEXT AND QUESTIONS** 

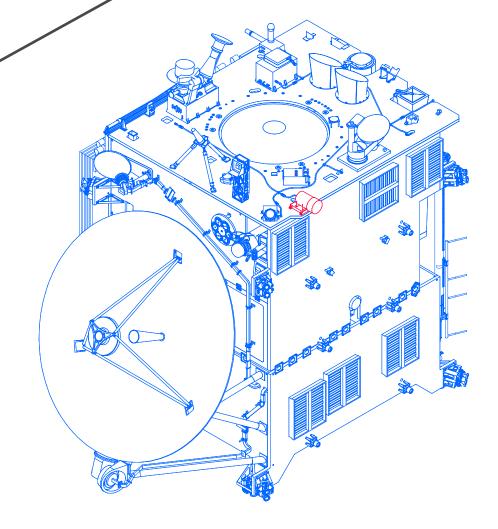
FAR FROM THE SUN - LOW ACTIVITY

In the study of the interaction between the solar wind and bodies of the solar system, there is one more parameter to consider for comets, compared to planets and moons: the distance to the Sun. The impact of this parameter on the solar wind dynamics around a nucleus is huge. We know that most of the comet present an atmosphere-less nucleus to the solar wind at aphelion, and an extended and dense coma at perihelion.

PERIHELION
A bow shock and a solar wind cavity are formed

APHELION
The solar wind directly impacts a bare nucleus

But what happens inbetween? How does a solar wind cavity form and grow, out of nothing, to a large and established cavity? For two years, the European Rosetta mission provided the ideal observation conditions for this precise topic.



## INSTRUMENT

The Ion Composition Analyzer Rosetta Plasma Consortium RPC-ICA: E: 10 eV - 40 keV FOV: 90x360°

Cadence:192 s

M: 1, 2, 4, 8, 16 and 32 amu/e

E<sub>convective</sub>

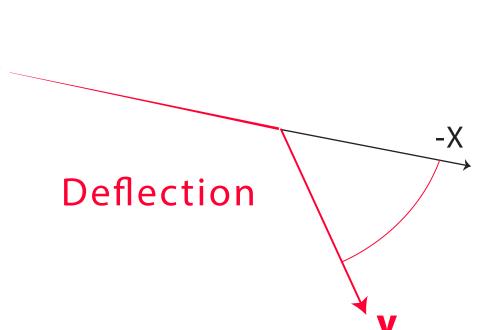
**SOLAR WIND** 

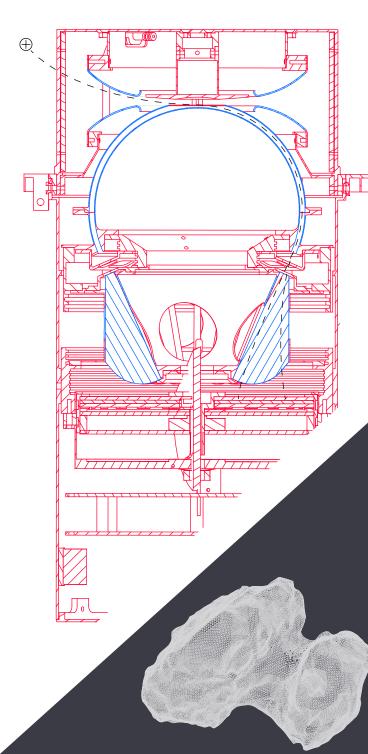
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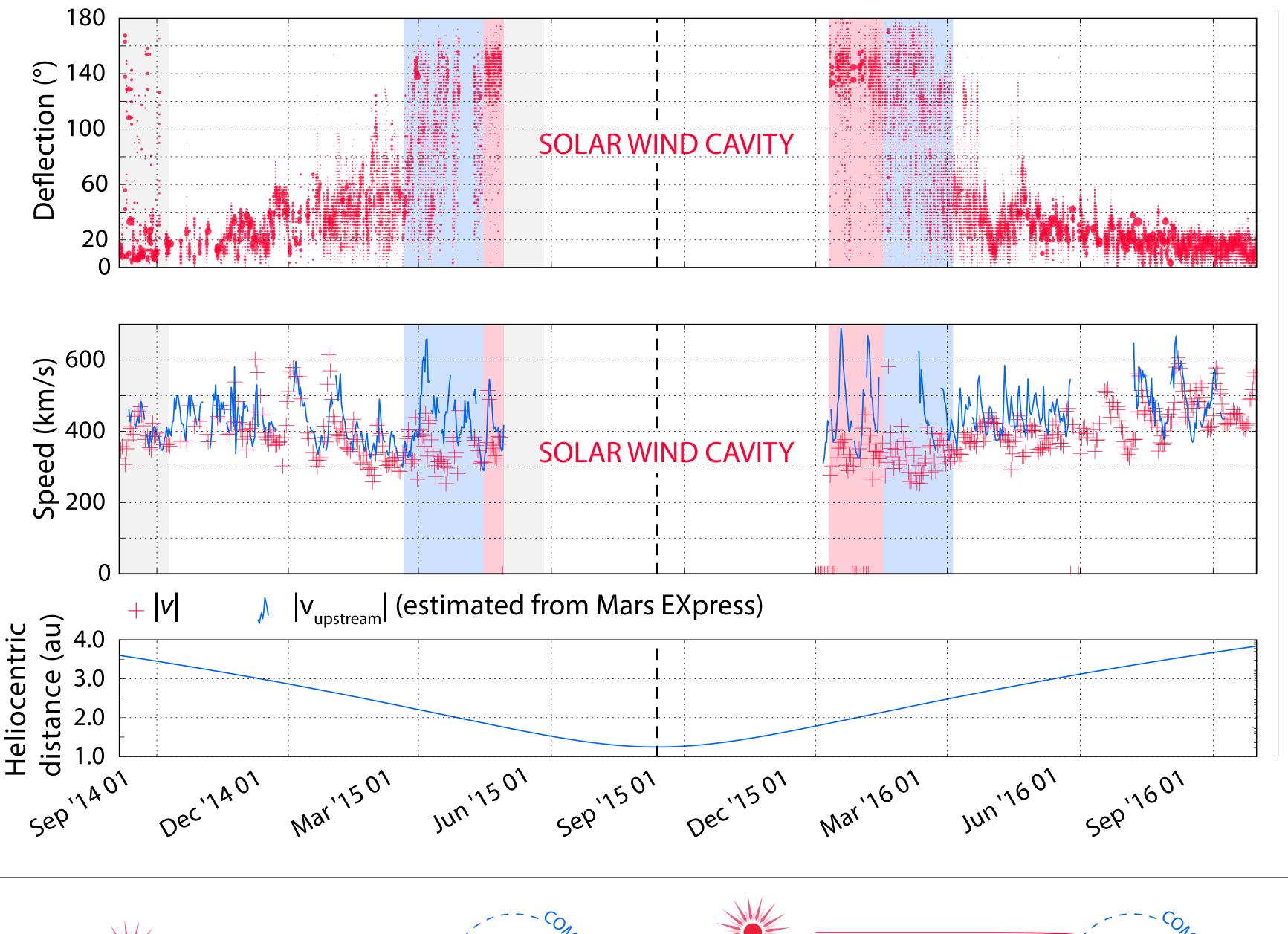
When sublimation and ionisation start: cometary ions are accelerated by the convective E-field, the solar wind loses the momentum and the enrgy gained by the cometary population. Two effects are expected on the solar wind:

- DEFLECTION

- DECELERATION







## OBSERVATIONS Proton bulk properties

- Deflection is in fact observed, but surpisingly up to 180°
- Clear solar wind cavity(>7 months, centered around perihelion)
- Deflection gets more and more spread, before refocusing at ~140° just before and after the cavity
- Little deceleration observed





90° deflection



>90° deflection

