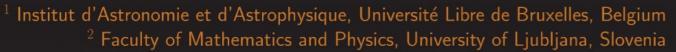
Characterization of SB1 detected in the Gaia-ESO Survey iDR5

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1. Data

- FLAMES/GIRAFFE HR10&21 spectra
- optical and IR wavelengths, R~20000 200000 single exposures from 50000 stars
- RV from recomputed CCFs (see Van der Swaelmen's poster)
- 4 observations per star with $\Delta T \leq 1$ week

2. Methods

- Statistical χ^2 -test per star with confidence level of 99.9% F2 distribution to assess the normality of RV uncertainties

3. Results

- SB1 sample contaminated by photometric variability

 → use of Gaia DR2 (parallaxes, G, BP, RP photometry) for:
 - discriminate dwarfs and giants
 - discriminate genuine SBIs from intrinsic RV vs. photometric dispersion

4. Characterization of SB1s

- ~700 SB1s/30000 stars: dwarfs 72% & giants 28%
- GES SB1 frequency: ~2%
 - per spectral type
 - per metallicity

