



THE DARK ENERGY SURVEY:

SURVEY STATUS AND SCIENCE RESULTS

EMMANUEL BERTIN

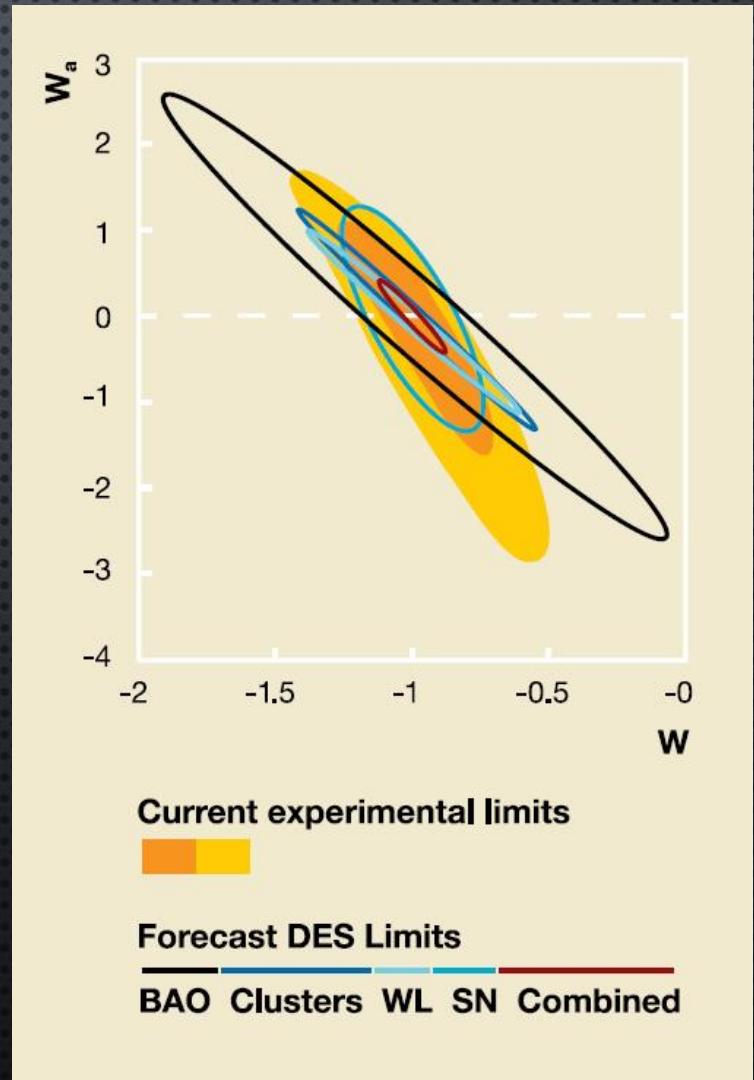
FOR THE DES COLLABORATION



DARK ENERGY
SURVEY

OVERVIEW

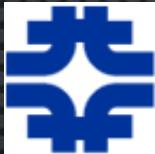
- 5000 SQ.DEGREE PHOTOMETRIC SURVEY OF THE SOUTHERN HEMISPHERE IN 5 BANDS (G,R,I,Z,Y) DOWN TO 24TH MAG (GALAXIES, 10σ)
 - 525 NIGHTS OVER 5 YEARS DURING 5-MONTH SEASONS
 - INCLUDES THE 2500 SQ.DEGREE SOUTH POLE TELESCOPE SPT-SZ SURVEY FOOTPRINT
 - 30 SQ.DEGREE REPEATED ~WEEKLY IN G,R,I,Z (SN FIELDS)
- SURVEY STARTED AUG 31, 2013.
- MAIN SCIENCE DRIVERS: 4 COSMOLOGICAL PROBES BASED ON DISTANCE, GEOMETRY AND STRUCTURE GROWTH:
 - GALAXY CLUSTER COUNTS ($\sim 100,000$) TO $z \sim 1$
 - GRAVITATIONAL LENSING (STRONG AND WEAK) FROM 200 MILLION GALAXIES
 - BARYON ACOUSTIC OSCILLATIONS FROM 300 MILLIONS GALAXIES TO $z \geq 1$
 - 4000 SUPERNOVA IA LIGHT CURVES WITH $0.1 < z < 1.1$





THE COLLABORATION

DARK ENERGY
SURVEY



- JOSH FRIEMAN, DIRECTOR
- PROJECT INITIATED IN 2004
- 28 INSTITUTIONS, ~400 PARTICIPANTS
- SCIENCE WORKING GROUPS: CLUSTERS, GALAXY EVOLUTION, LARGE SCALE STRUCTURE, MILKY WAY, PHOTOMETRIC REDSHIFTS, QSOs, STRONG LENSING, SUPERNOVAE, THEORY, WEAK LENSING
- FUNDING: DOE, NSF, DFG, CSIC, CNPQ, FAPERJ, FINEP + INSTITUTIONS
- NO FRENCH INSTITUTION INVOLVED
 - BUT A FEW FRENCH INDIVIDUALS INVOLVED AS PHDs, POST-DOCS OR EXTERNAL COLLABORATORS
 - POSSIBILITY TO SUBMIT EXTERNAL COLLABORATOR PROPOSALS

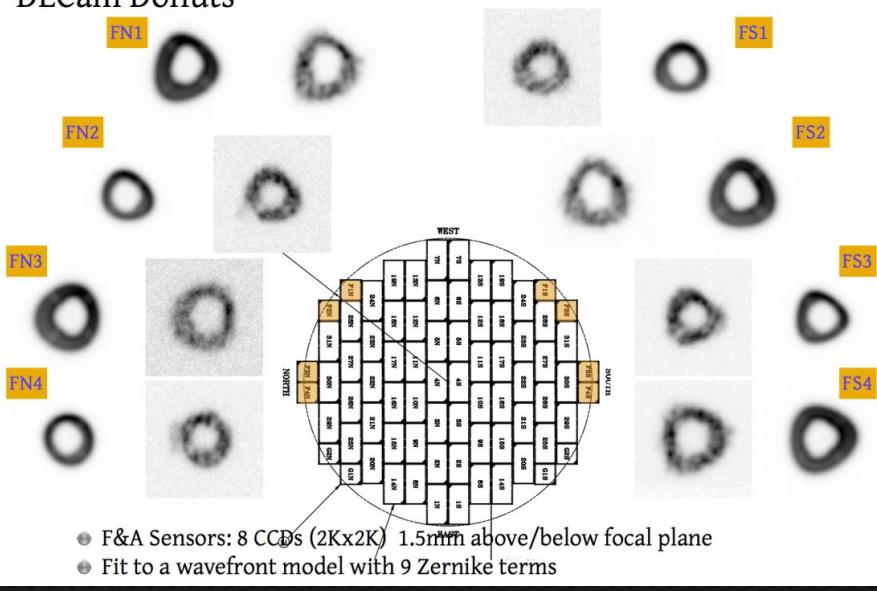


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THE INSTRUMENT: DECAM

- INSTALLED ON THE BLANCO 4M TELESCOPE AT CTIO. SEEING (I BAND) $\sim 0.9''$
- 3-SQ. DEGREE FIELD OF VIEW (2.2 DEGREE DIAMETER).
- HEXAPOD COMPENSATES FLEXURES BASED ON OUT-OF-FOCUS IMAGES

DECam Donuts

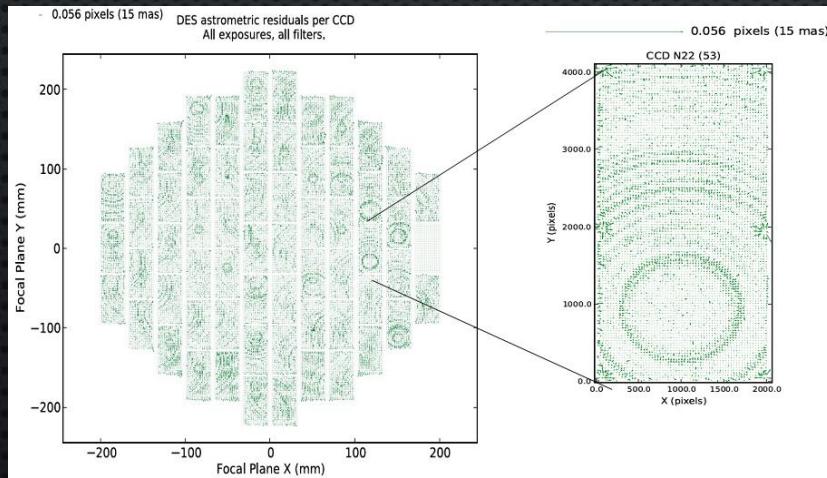
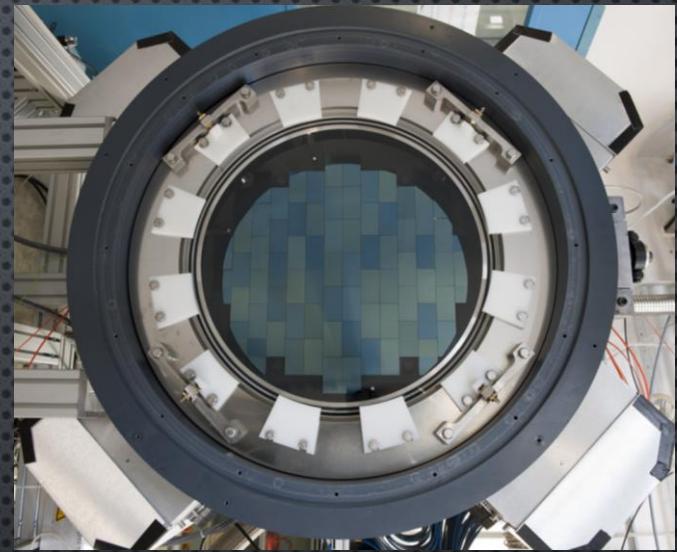




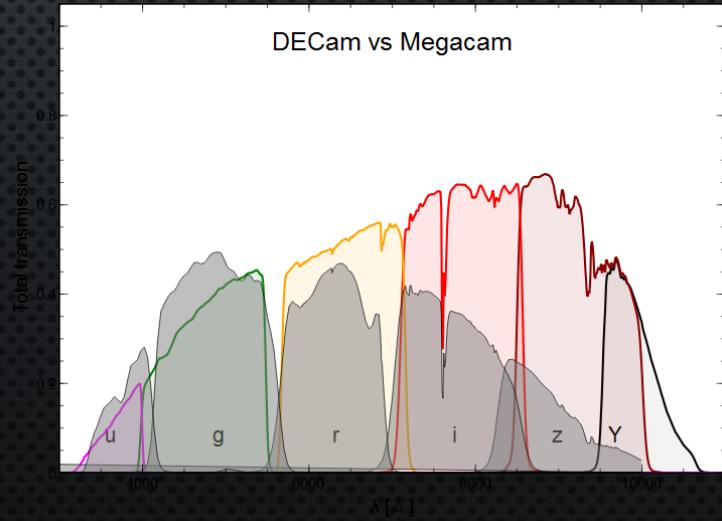
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DECAM DETECTORS

- 570 MPIXELS
- THICK, FULLY-DEPLETED CCDs
 - VERY HIGH Q.E IN THE RED
 - VERY LITTLE FRINGING IN I AND Z
 - CARE SHOULD BE TAKEN WITH STRONG LIGHTING (SUPER-SATURATION)
 - SIGNIFICANT DISTORTION OF THE PIXEL GRID DUE TO LATERAL ELECTRIC FIELD VARIATIONS
 - STATIC: RESISTIVITY VARIATIONS ("TREE-RINGS")
 - DYNAMIC: "BRIGHTER-FATTER" EFFECT



A. Plazas

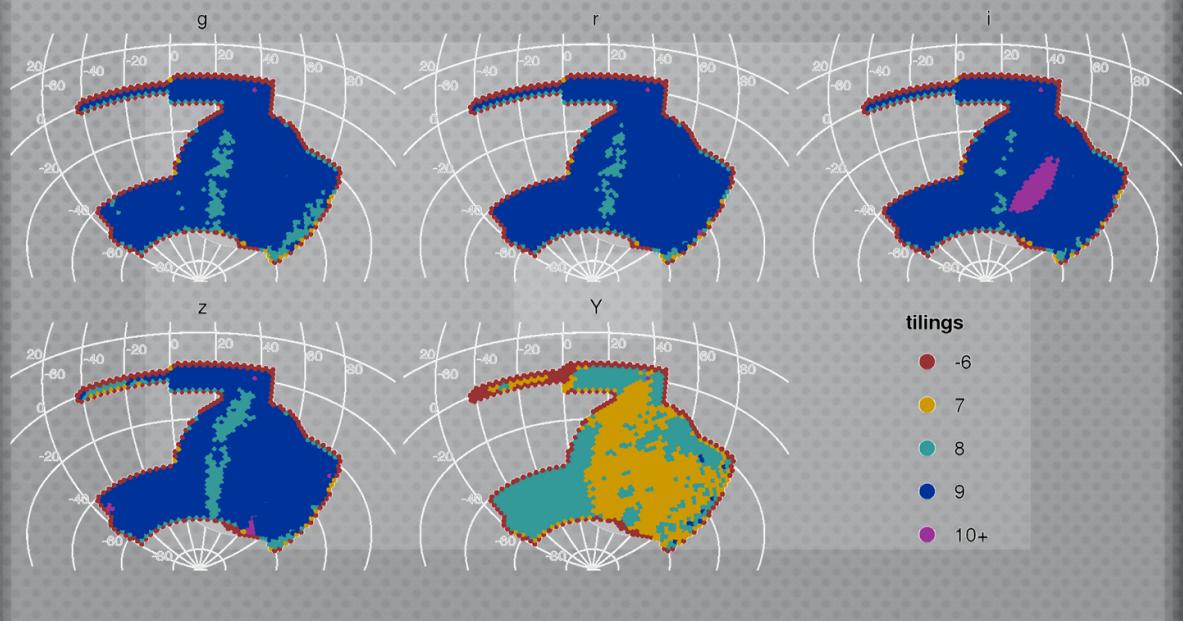




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SURVEY AND OBSERVATION STRATEGY

- OBSERVATIONS ARE CARRIED OUT BY DES MEMBERS FROM AUGUST TO FEB
- ~ 90s EXPOSURES
- 12 EPOCHS, ~2 PER SEASON
- SWITCH TO SN FIELDS DEPENDING ON OBSERVING CONDITIONS
- VERY LARGE DITHERS (UP TO 1 DEG) IMPLY
 - MULTI-EPOCH MEASUREMENTS



E.Nielsen

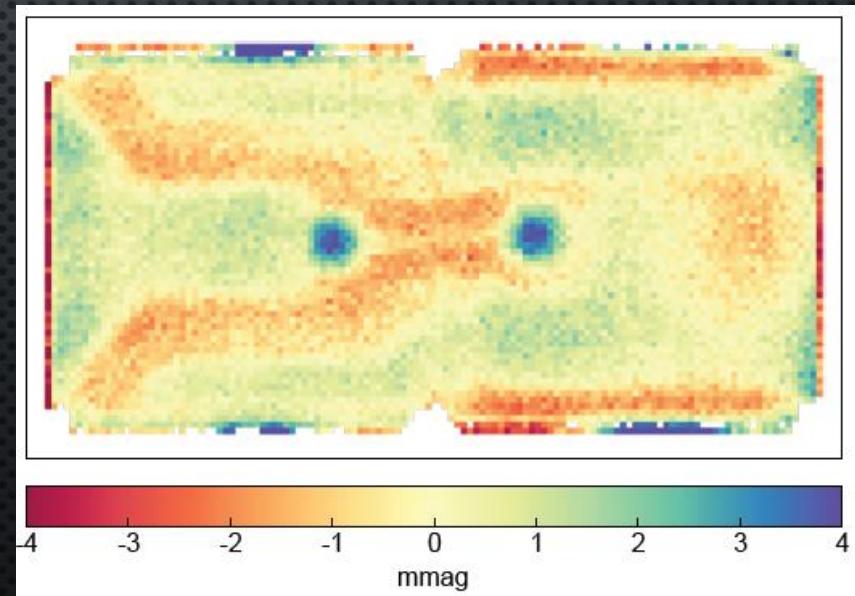
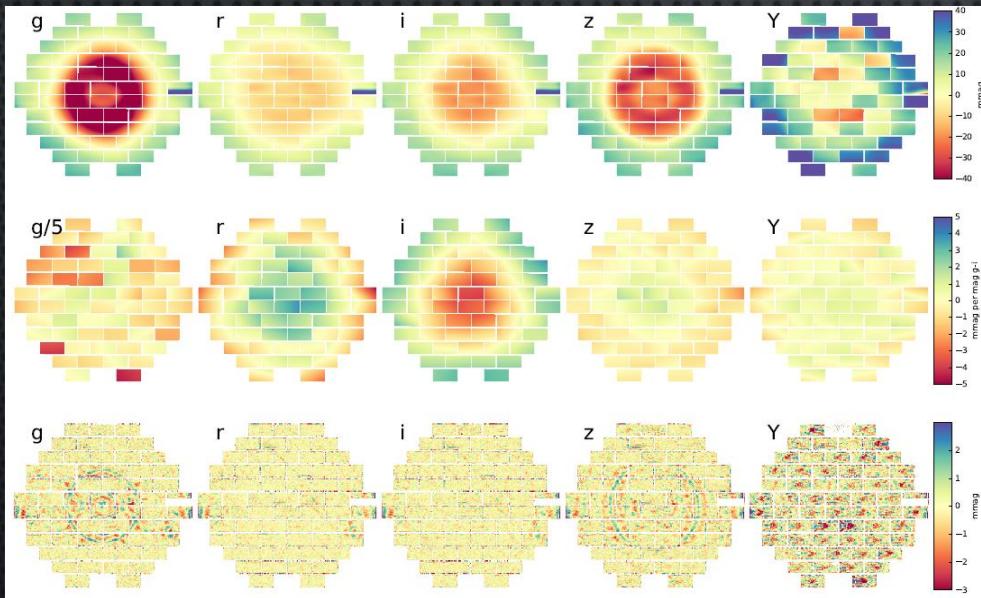




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INSTRUMENTAL CALIBRATION

- PHOTOMETRIC RESPONSE MODELLED AT THE MILLIMAG LEVEL (PHOTOMETRIC HOMOGENEITY REQUIREMENT: <20 MILLIMAG).
- ASTROMETRIC RESPONSE MODELLED AT THE MAS LEVEL (RELATIVE ASTROMETRY REQUIREMENT: <15 MAS)



G.Bernstein



DATA

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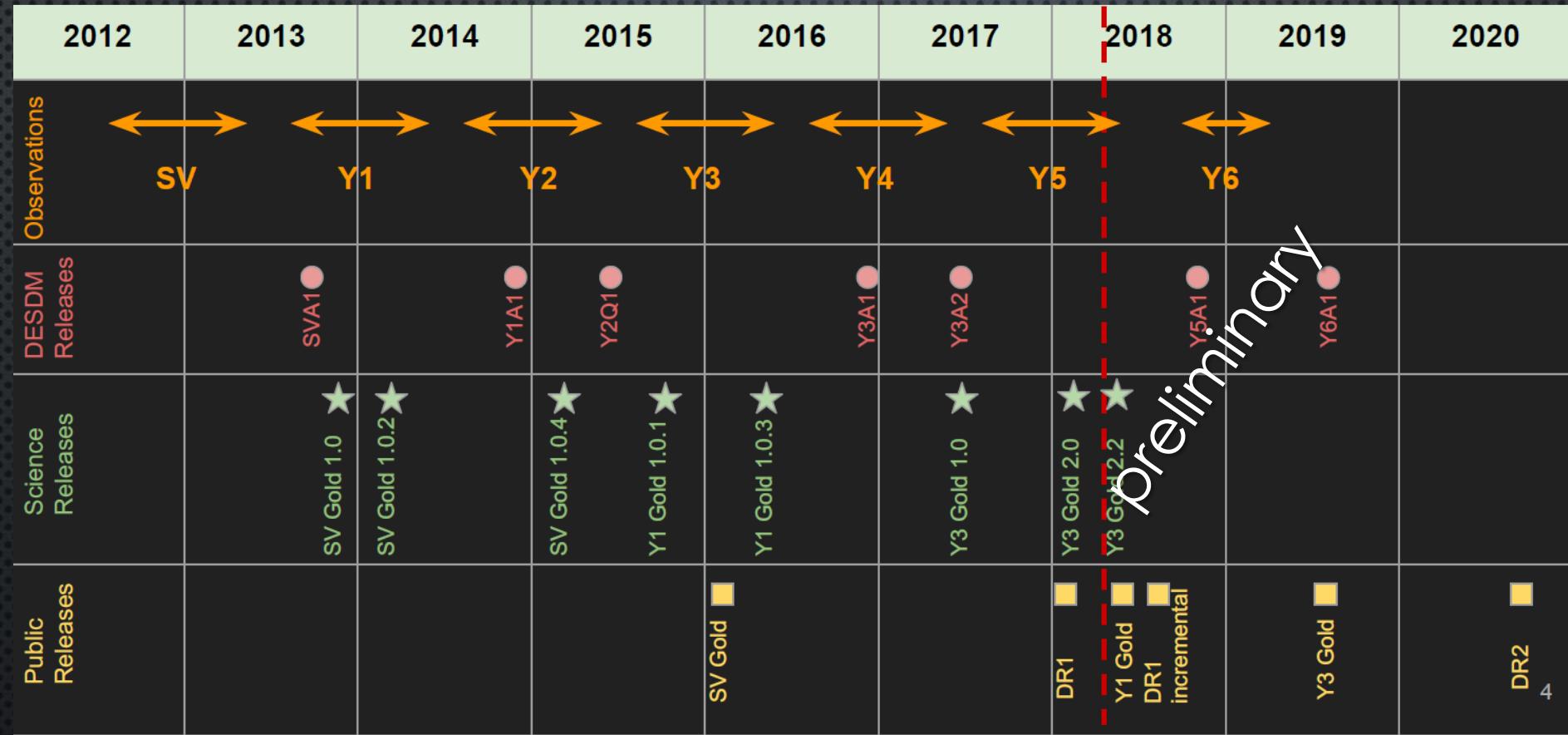
- SCIENCE DATA THROUGHPUT ~400GB PER CLEAR NIGHT
- DATA PROCESSED AT NCSA (U. OF ILLINOIS) THROUGH THE DESDM (DATA MANAGEMENT) SYSTEM
 - DECam COMMUNITY PIPELINE OPERATED BY NOAO
 - DES DATA PRODUCTS READILY AVAILABLE TO THE DES COMMUNITY THROUGH THE DES ARCHIVE AT NCSA
- YEARLY / BI-YEARLY INTERNAL DATA RELEASES TO THE COLLABORATION
- RAW DATA PROPRIETARY PERIOD IS 12 MONTHS (NOAO STANDARD: 18 MONTHS)
 - CAN BE ACCESSED THROUGH THE NOAO NVO PORTAL





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DATA RELEASES



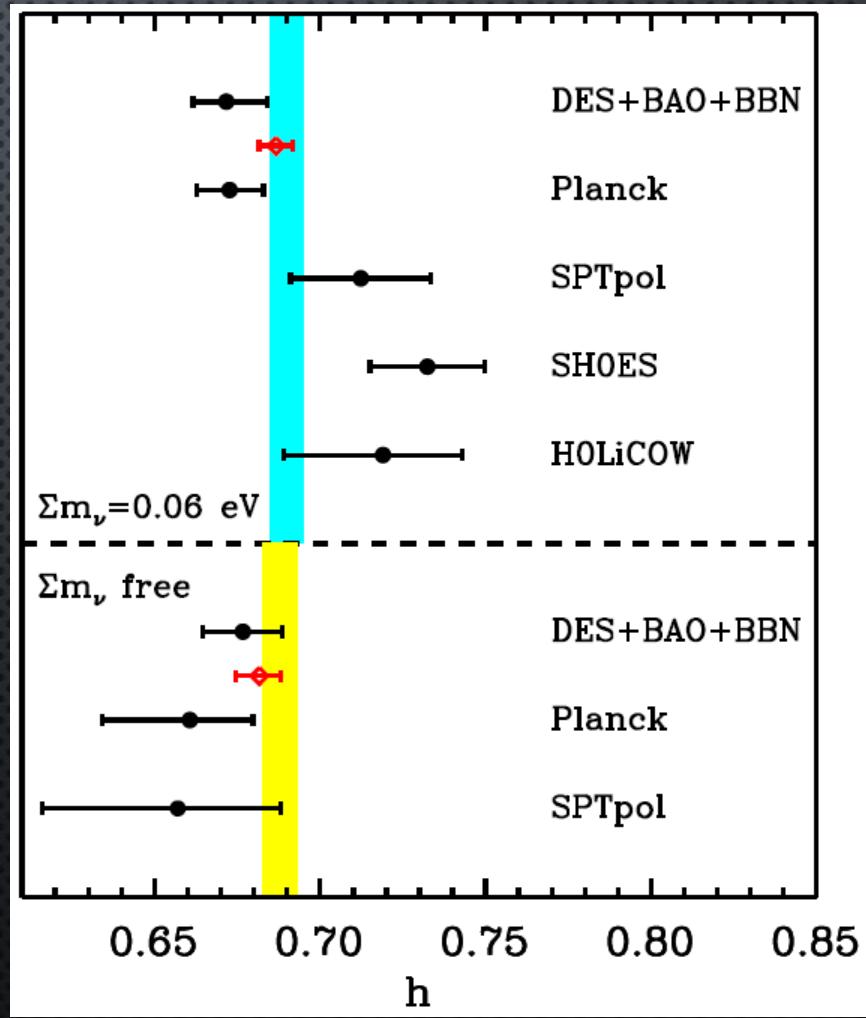
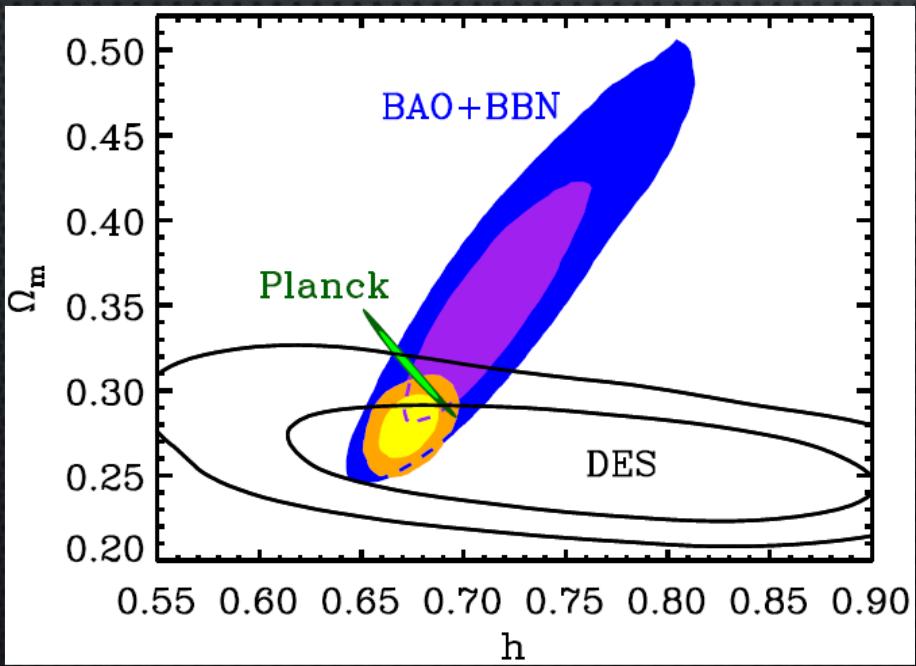
Courtesy of M. Carrasco-Kind



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YEAR 1 SCIENCE HIGHLIGHTS

- BAO+BBN CONSTRAINS: $H_0 = 67.2 \pm 1 \text{ km.s}^{-1}$
(ABBOTT ET AL. 2017)

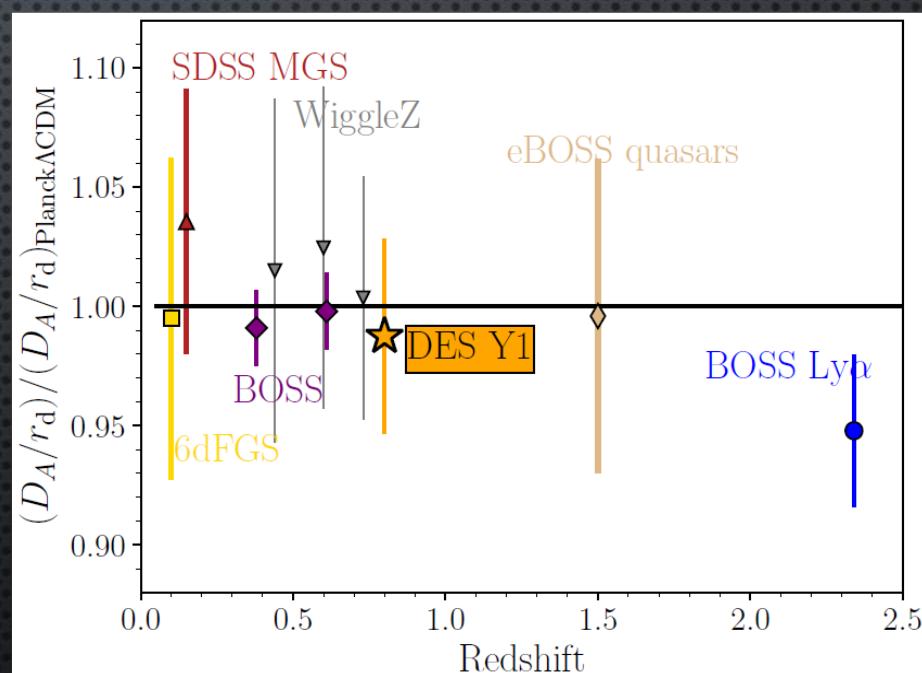
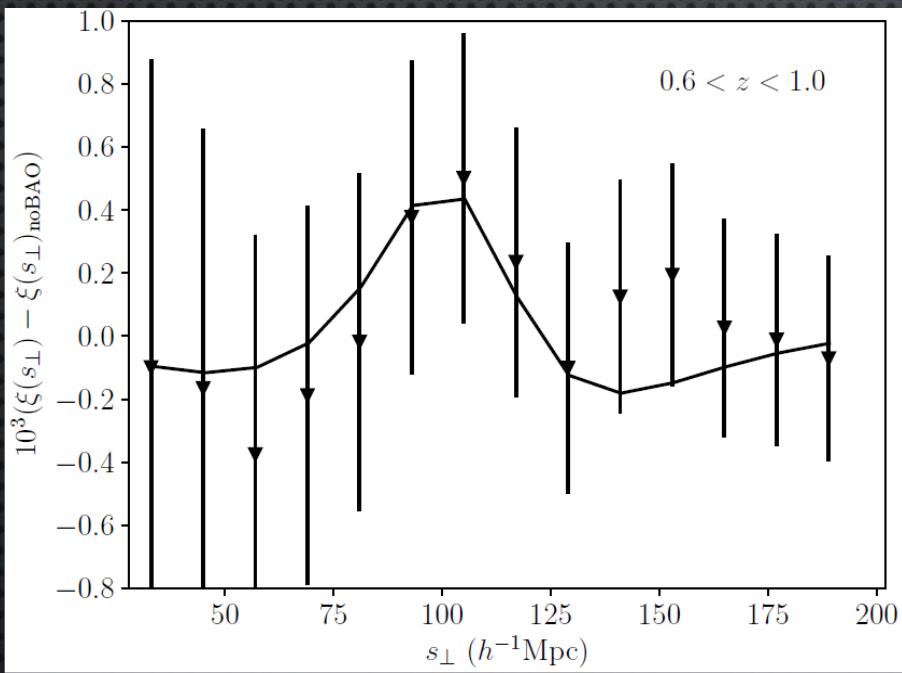




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YEAR 1 SCIENCE HIGHLIGHTS

- BAOs (ABBOTT ET AL. 2017)

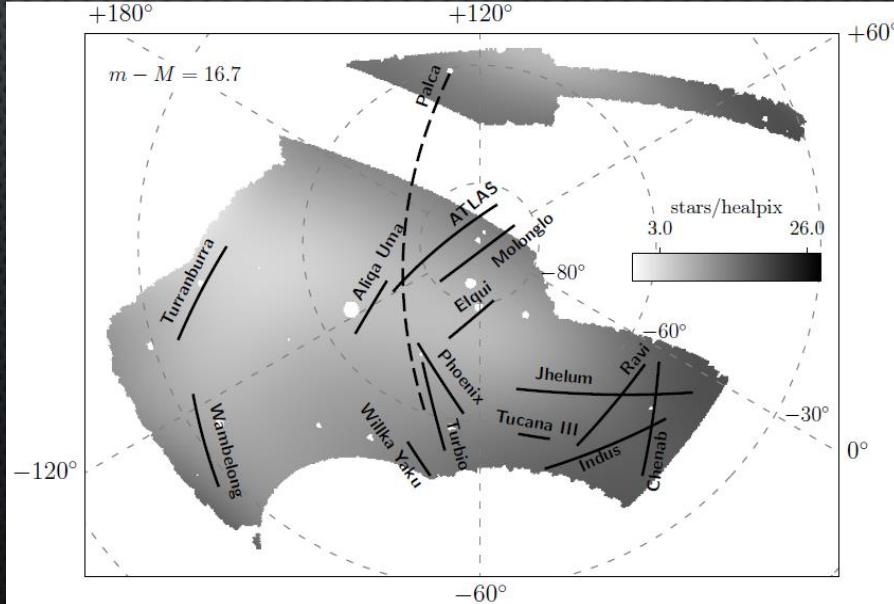
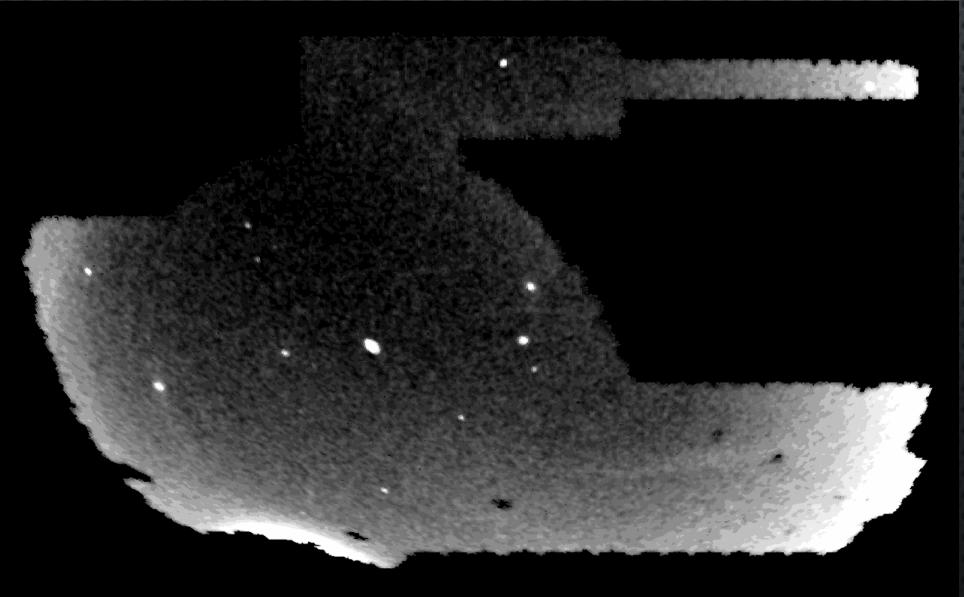
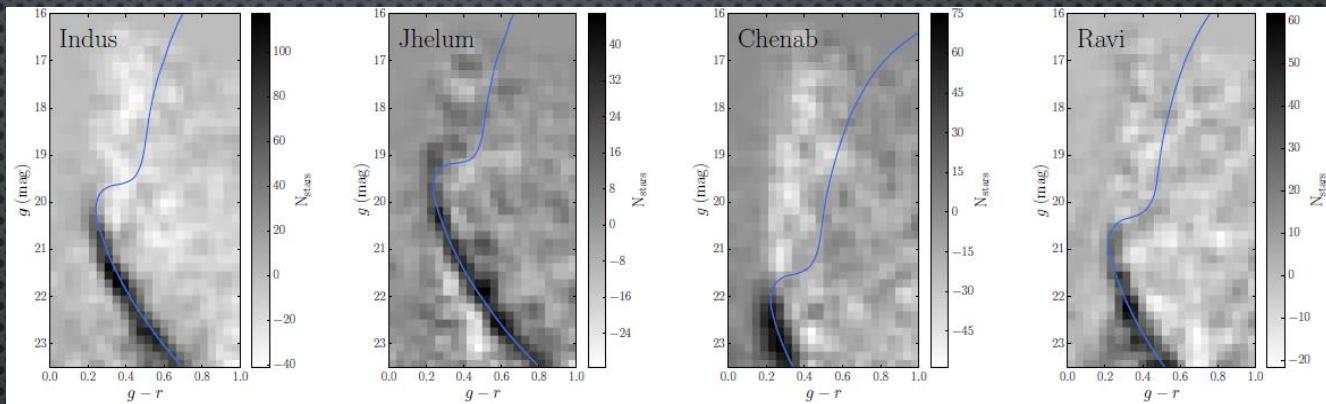




EARLY YEAR 3 RESULTS

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- STELLAR STREAMS
(SHIPP ET AL. 2018)
- 11 NEW STREAMS

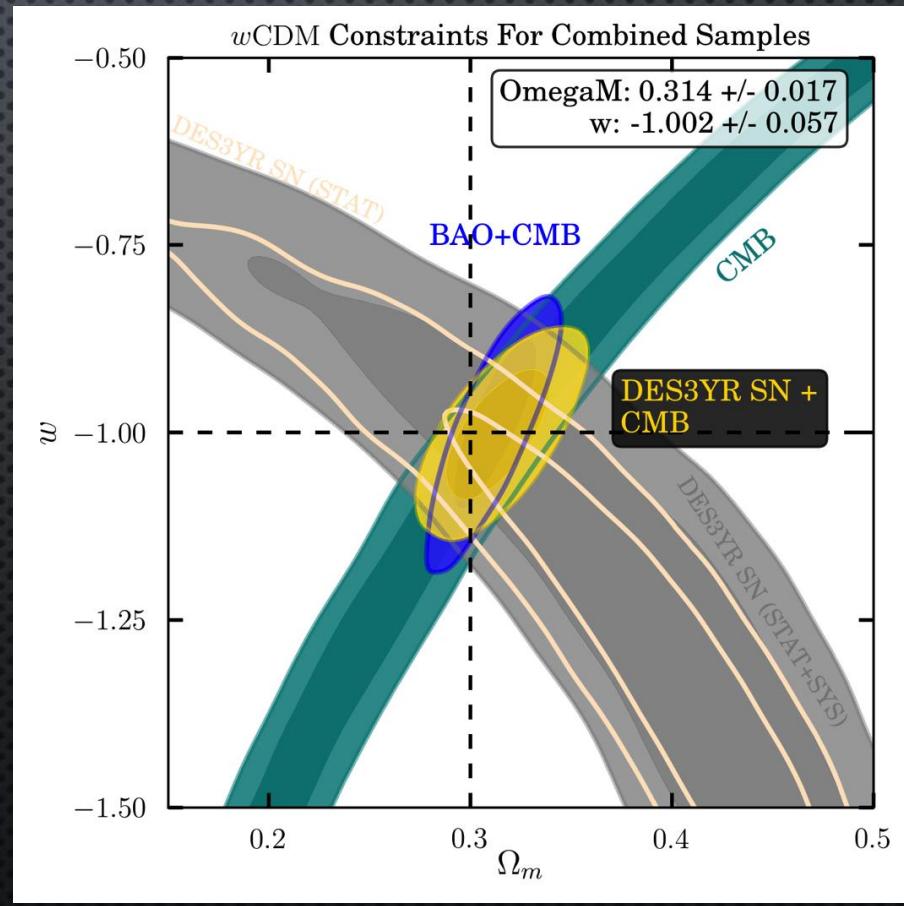
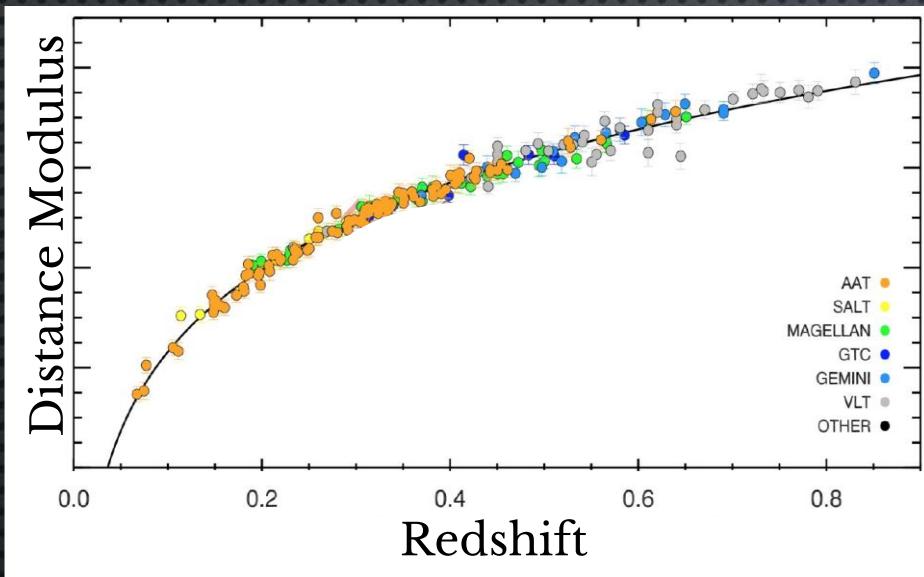




EARLY YEAR 3 RESULTS

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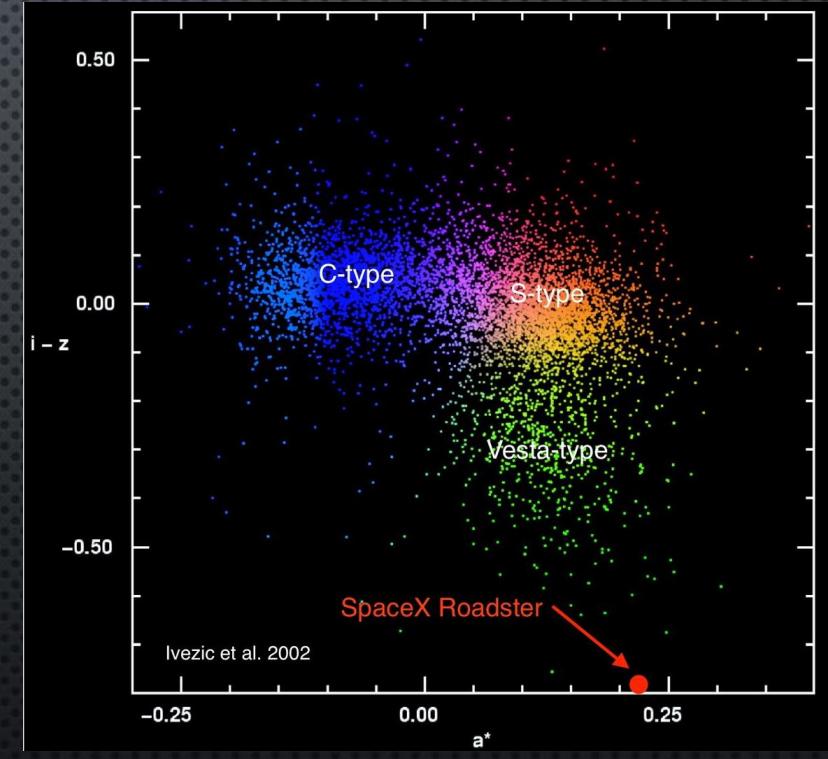
- BROUT ET AL. (AAS 2018)
- PRELIMINARY





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TRANSIENT AND MOVING OBJECTS



D.Guerdes