Periodicity analysis of young stars in Taurus

NOEMI ROGGERO

IPAG – GRENOBLE

1ST YEAR PHD STUDENT

SUPERVISORS: J. BOUVIER, F. MENARD







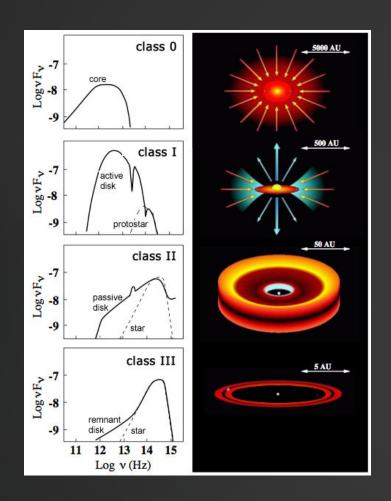


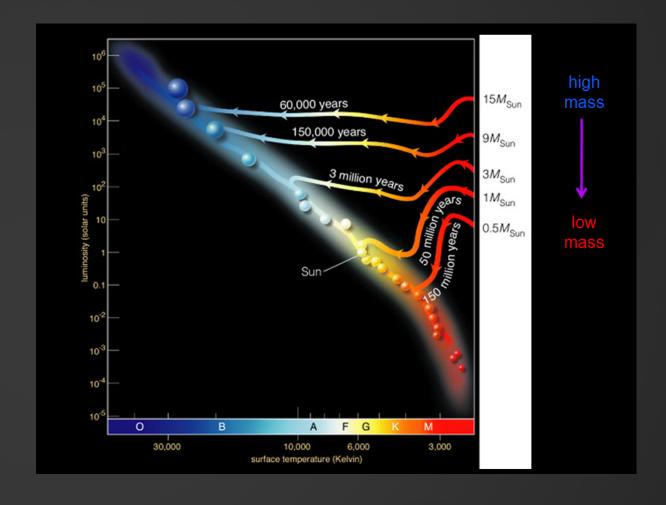


Photo: NASA

- ▶Introduction: what are dippers?
- ▶ Wavelet analysis
- ▶ Results
- ▶ Outlook

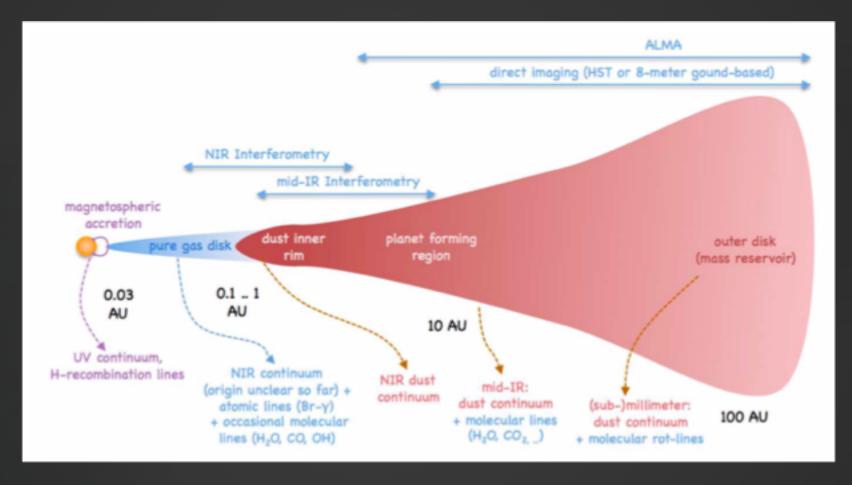
Low-mass star formation





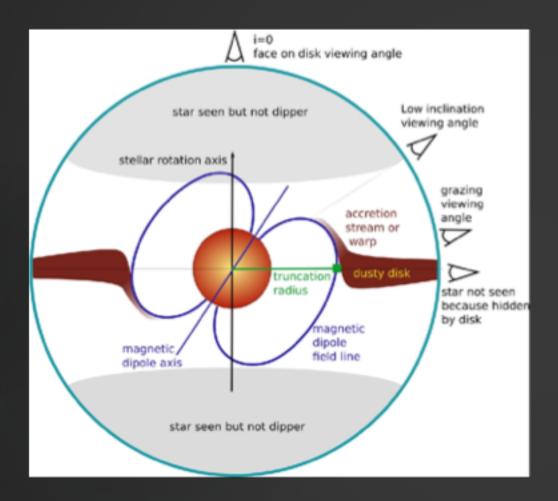
A. Isella 2006

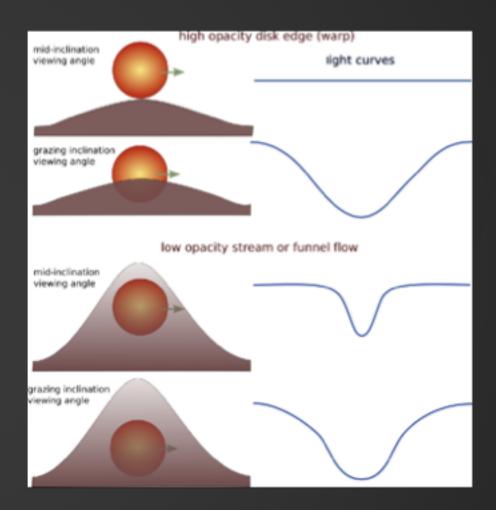
Protoplanetary disks



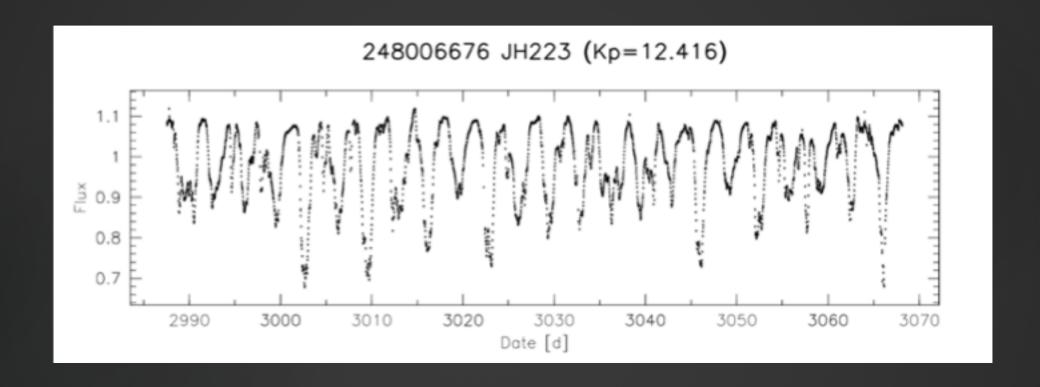
Dullemond & Monnier 2010

What causes dippers?





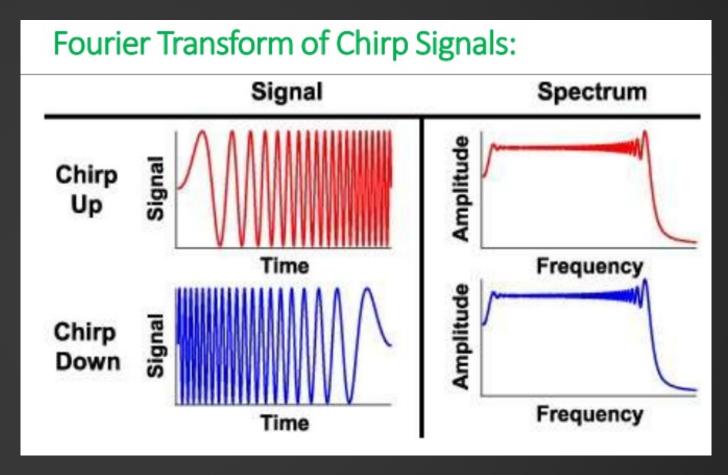
A prototypical dipper: JH223



- ▶ Introduction: what are dippers?
- ▶ Wavelet analysis
- ▶ Results
- ▶ Outlook

The Fourier transform

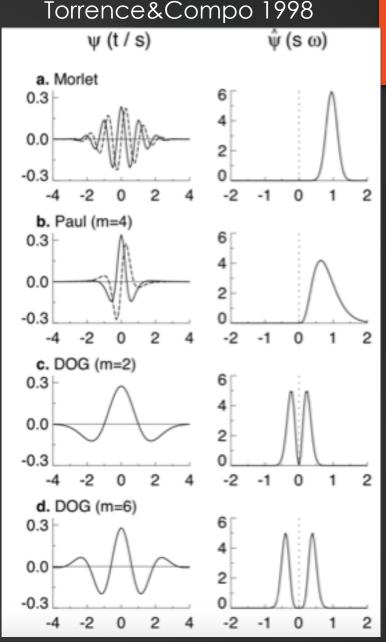
- ► Finds frequencies in the time series with high frequency resolution
- Problem: no time resolution for transient frequencies



Source: Berkouk & Sadmi 2018

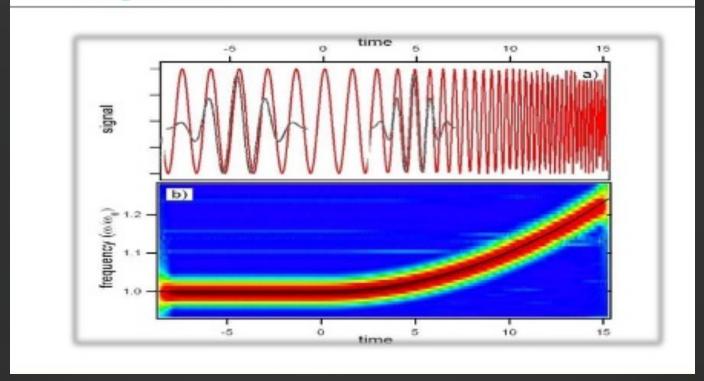
Wavelet transform

▶ Idea: consider a finite wave (instead of an infinite sinus) and shift it along the time series. The transform is the convolution of the signal and the wavelet at each position



Example: wavelet transform of the chirp signal

Time-frequency representation of « up-chirp » signal using CWT:

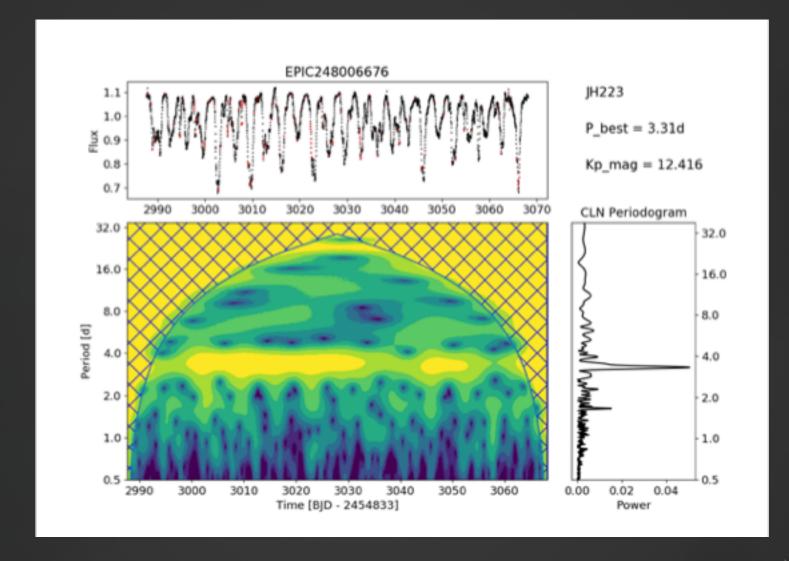


- ▶ Introduction: what are dippers?
- ▶ Wavelet analysis
- **▶** Results
- ▶ Outlook

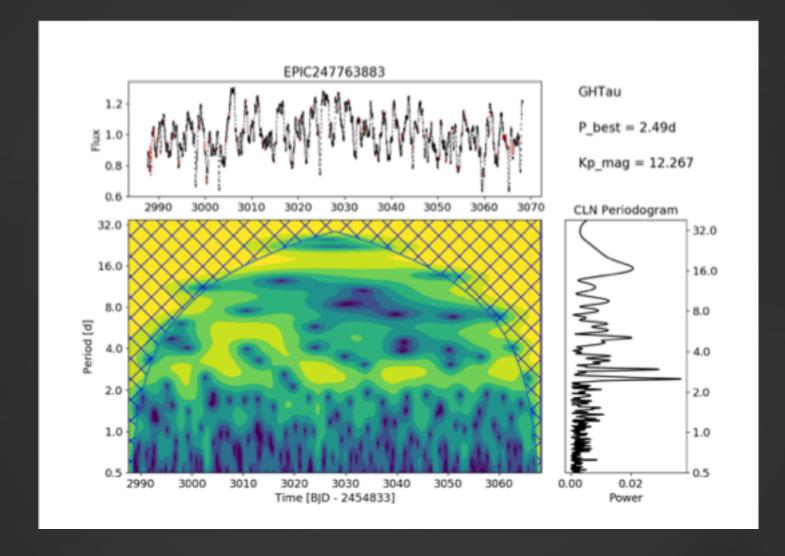
Data set

- ▶ 230 accreting YSOs in Taurus are observed in the K2 campaign (2017)
- ▶ Duration of the campaign: ~3 months, cadence 30min
- ~35 objects are dippers or show the dipper phenomenon together with another variability (bursts, spots etc.)
- ▶ In previous studies, 20-30% of CTTSs were classified as dippers (Cody et al. 2014, Alencar et al. 2010)

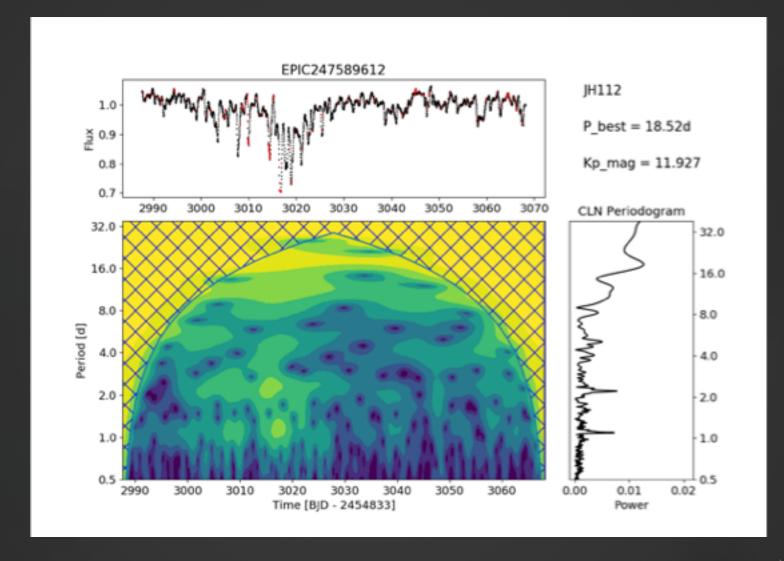
JH223: a dipper example



The case of multiple periodicities

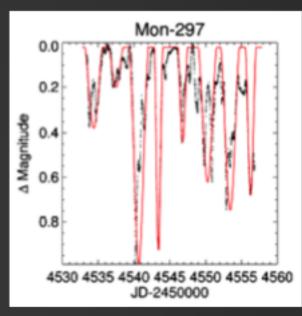


Transient dippers



Outlook

- ▶ It is possible to fit dippers with a simple occultation model
- ► Aim of the project: full radiative transfer modeling with the MCFOST code (Pinte et al. 2006)
- Search for planet signatures with simulations



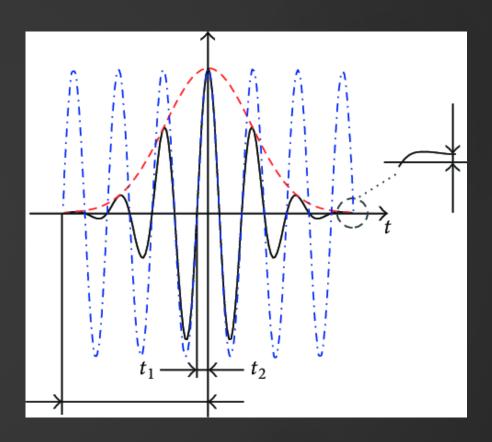
McGinnis et al. 2015

Morlet Wavelet

$$\psi(x) = \frac{1}{\sqrt{\pi T_b}} e^{-\frac{x^2}{T_b}} e^{2\pi i F_c x}$$

▶ Gaussian with width defined by T_b + complex wave

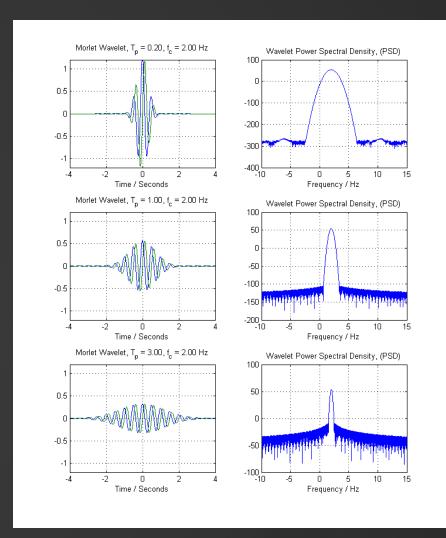
► To investigate all frequencies, the wavelet is stretched according to a scale s



Time vs. frequency resolution

- Increasing the width of the Gaussian increases the frequency and worsens the time resolution
- -> uncertainty principle
- Remember:

$$\hat{\psi}(\omega) = e^{-\frac{\pi^2(f - F_c)^2 T_b}{2}}$$



AA Tau

