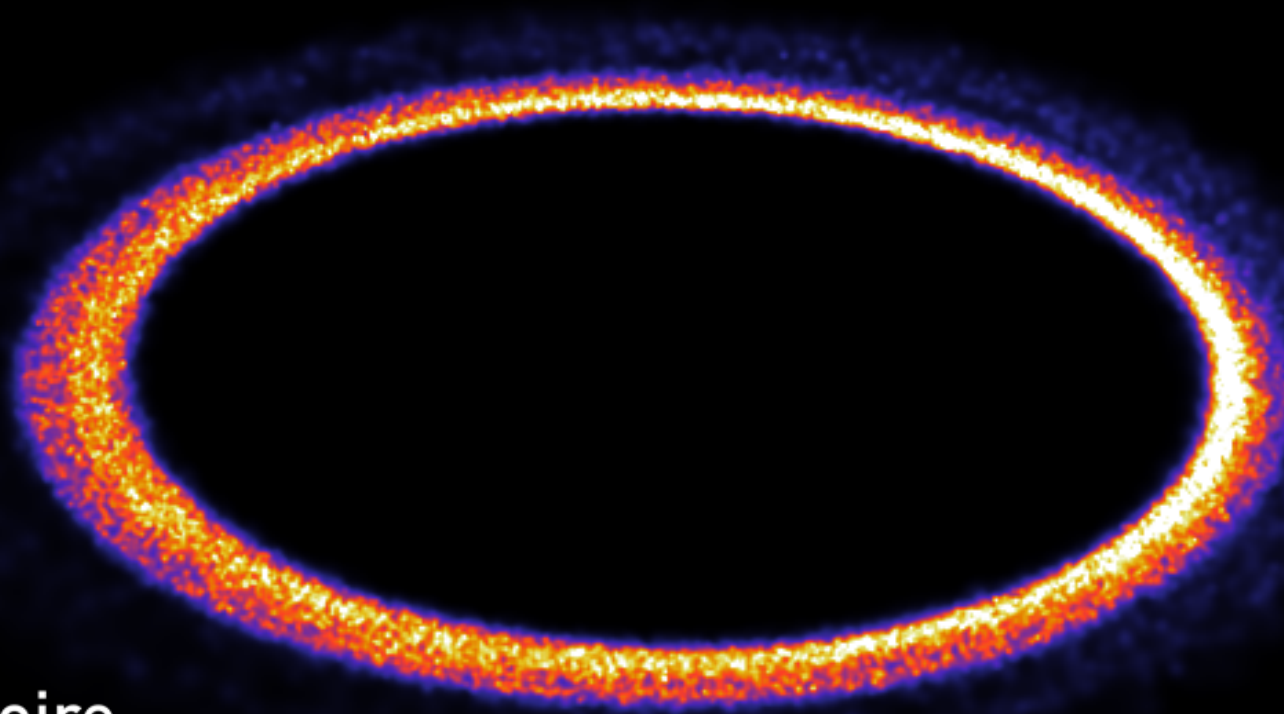


SELF-INDUCED DUST TRAPS AROUND SNOWLINES IN PROTOPLANETARY DISCS



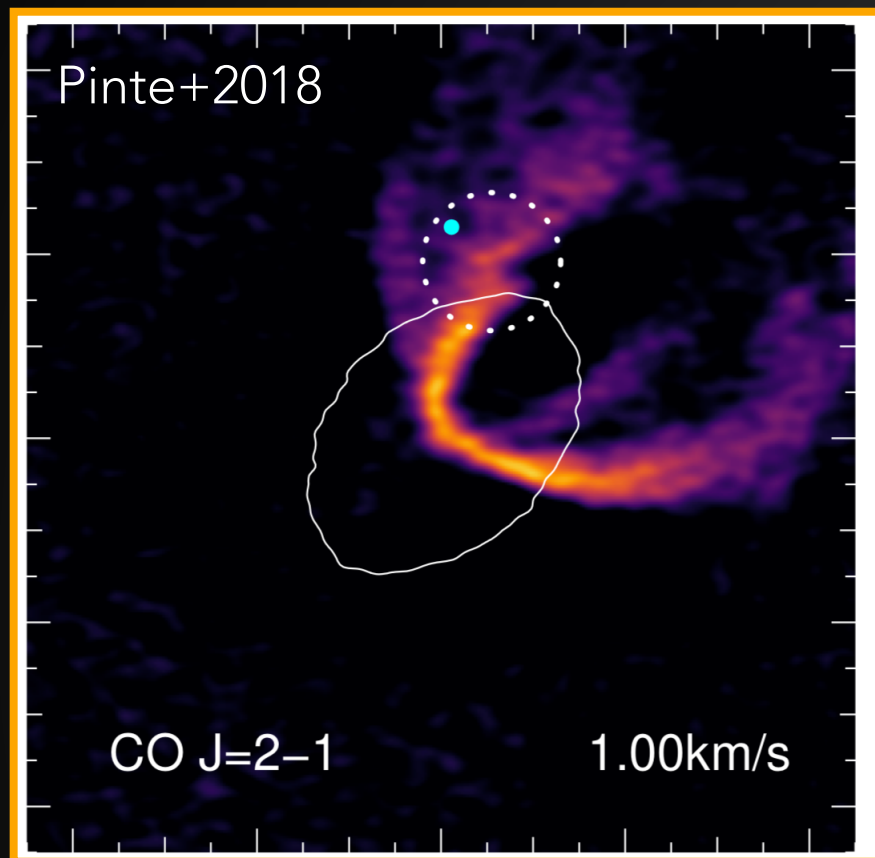
SF2A 2019, Nice



Arnaud Vericel, Jean-François Gonzalez

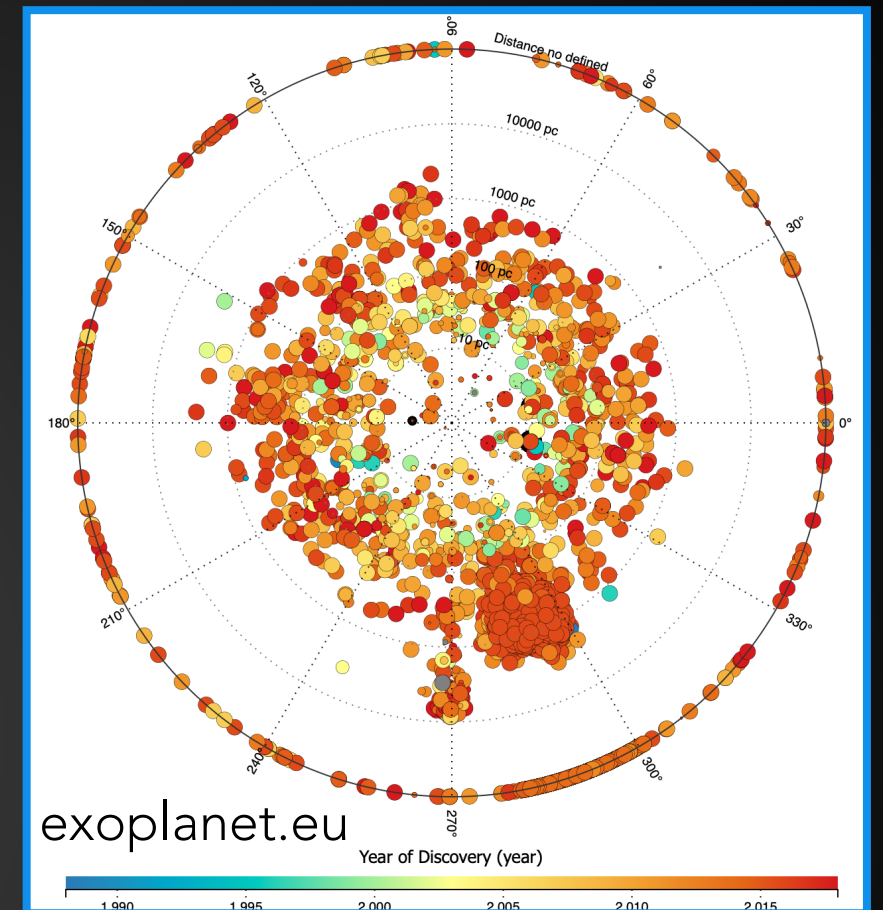
Centre de Recherche Astrophysique de Lyon

WHY DO WE CARE ?

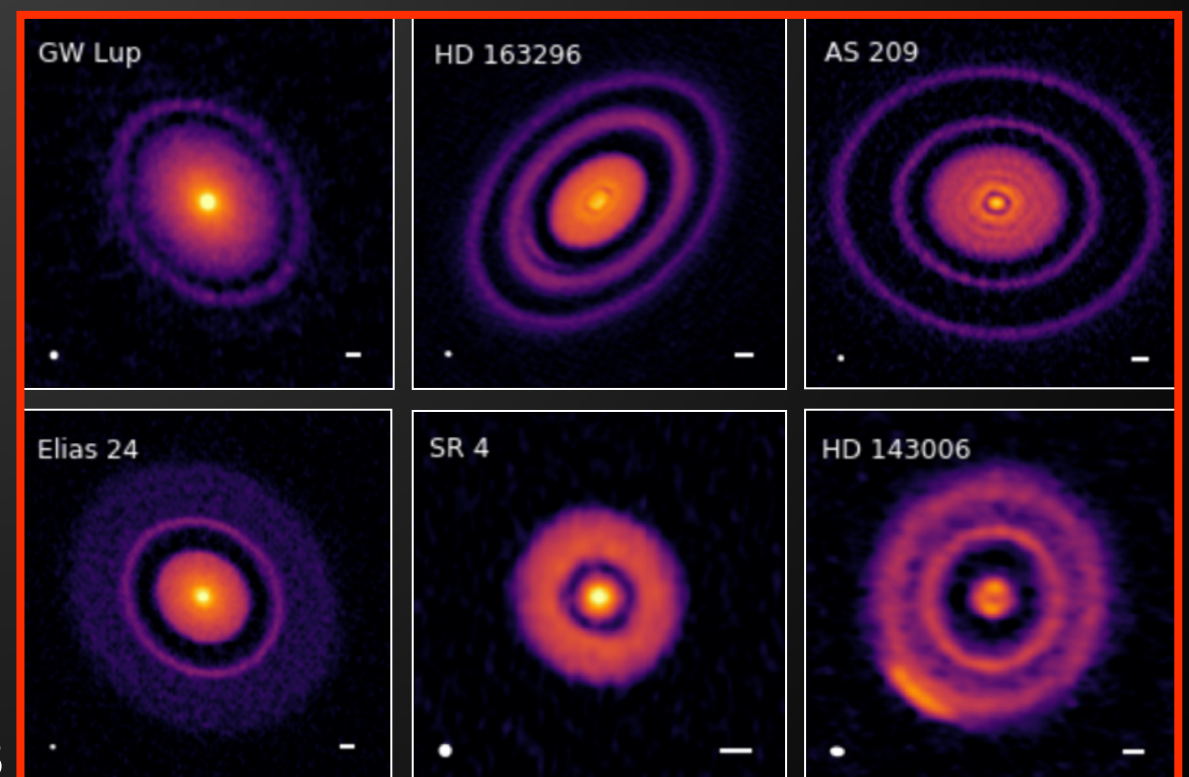


Planets during
protoplanetary disc phase

Exoplanet population
growing
and
diversified

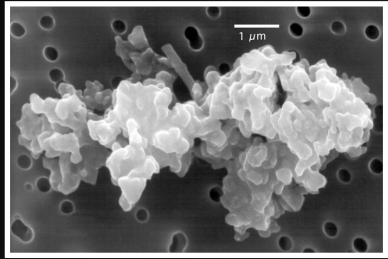


Dust structures in the mm
with ALMA



Andrews+2018

... BUT!

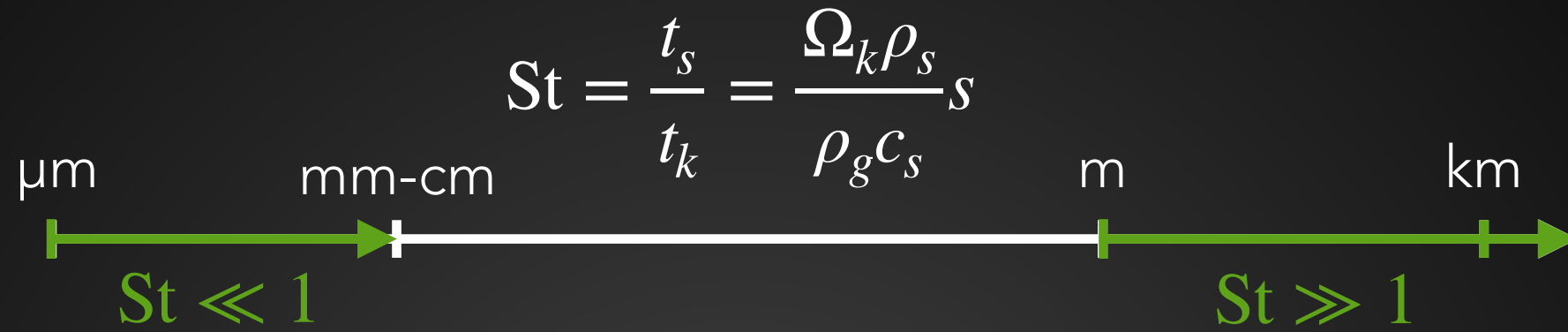
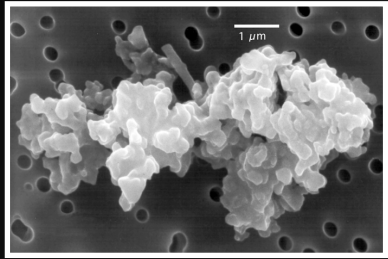


$$St = \frac{t_s}{t_k} = \frac{\Omega_k \rho_s}{\rho_g c_s} s$$

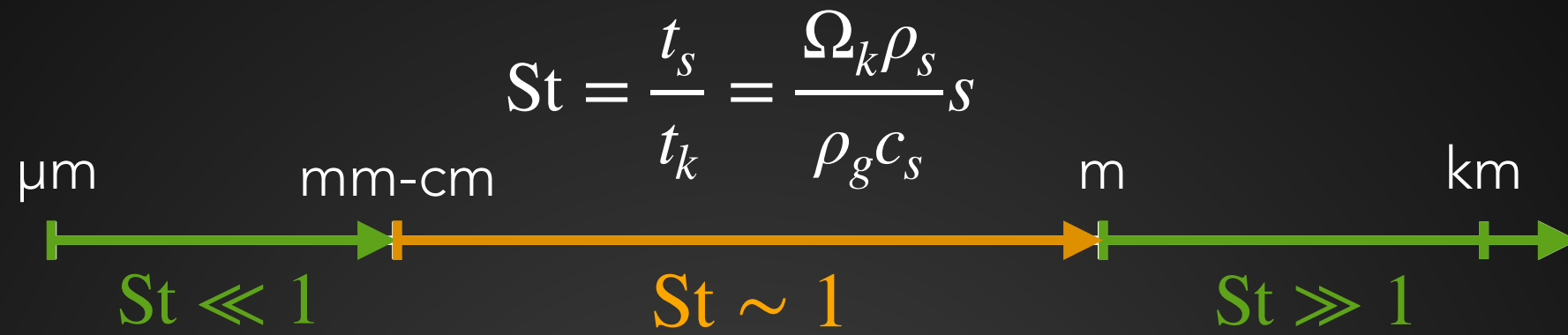
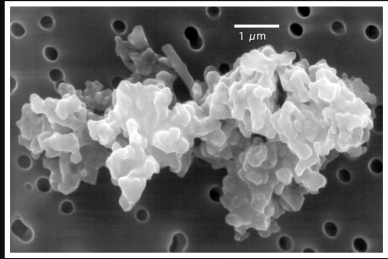
μm mm-cm m km



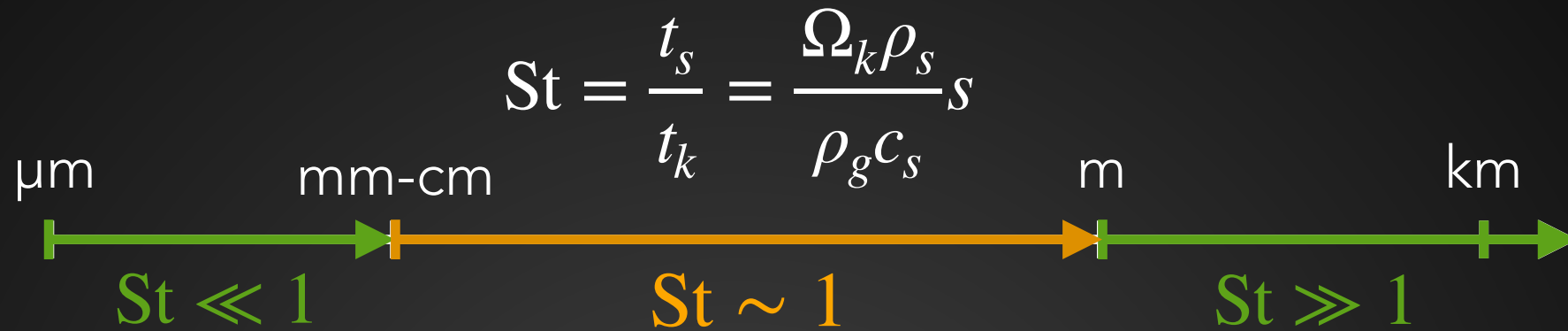
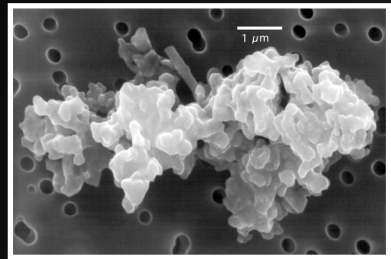
...BUT!



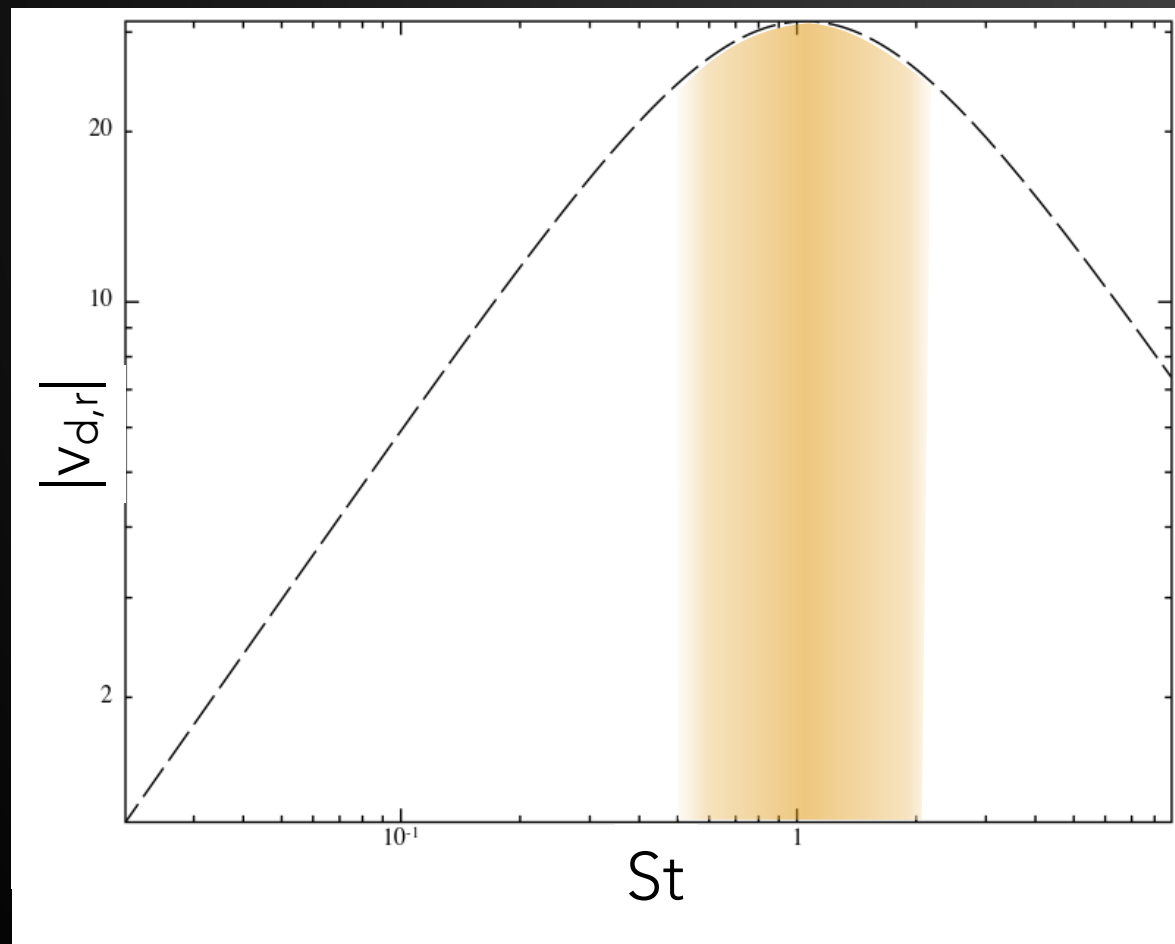
... BUT!



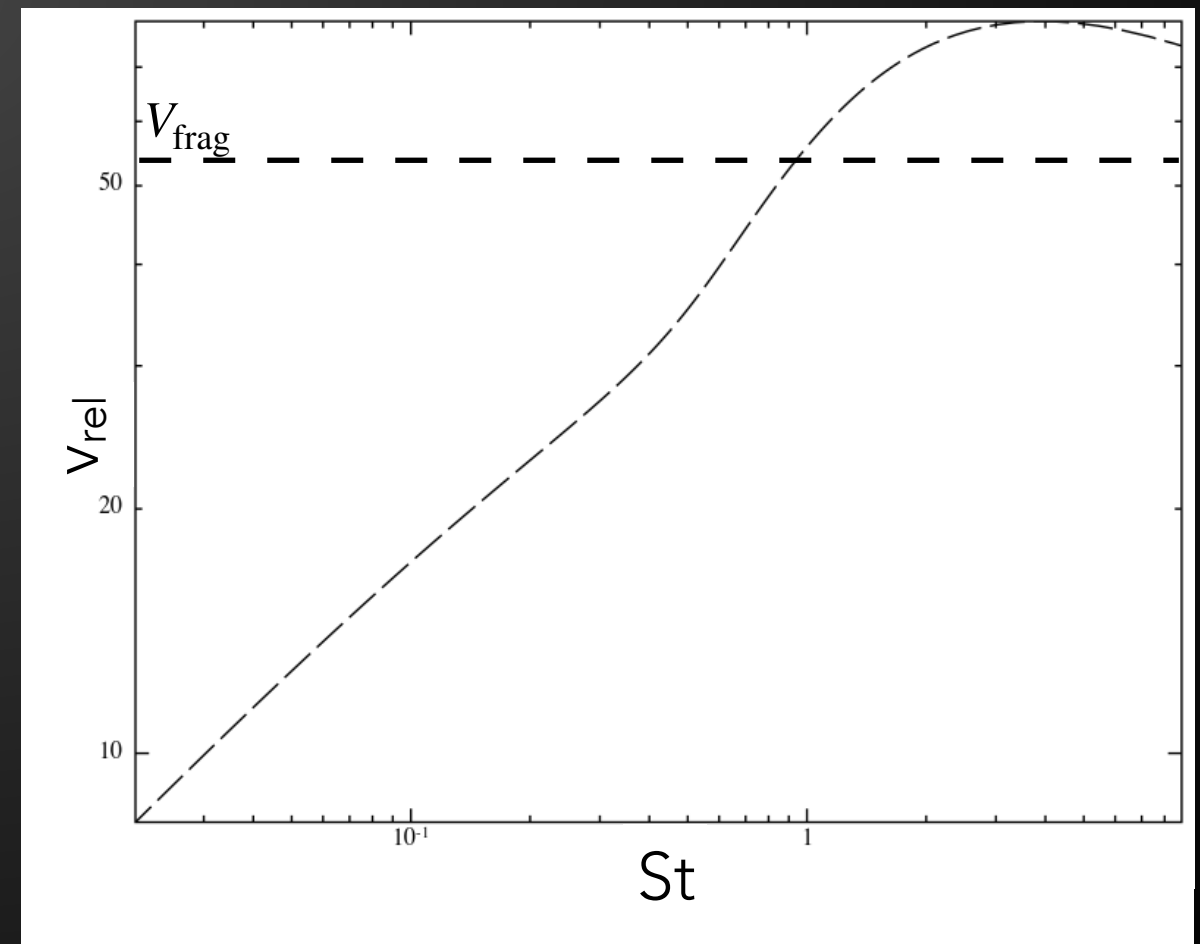
...BUT!



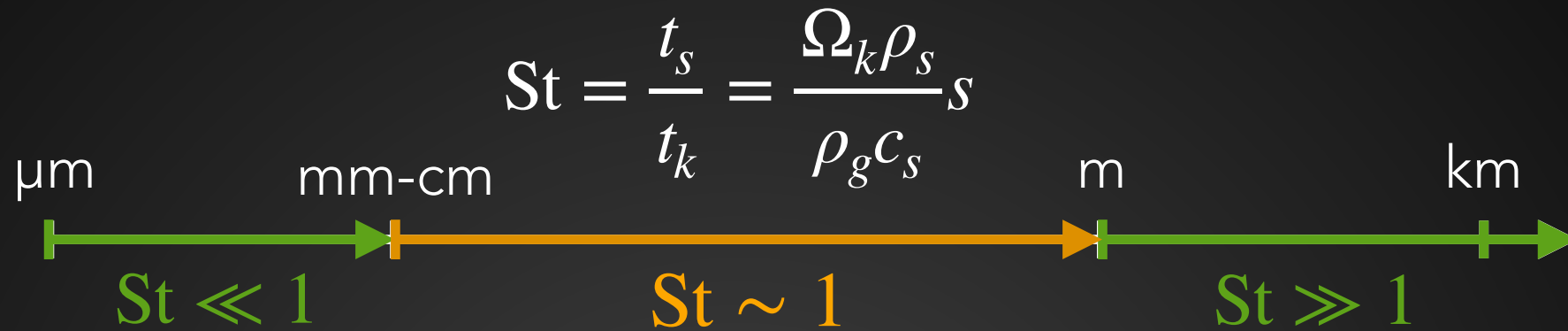
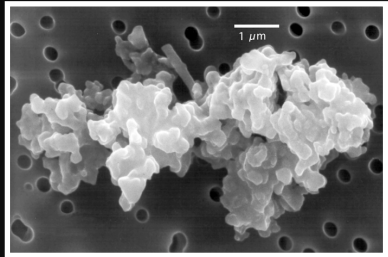
“Radial drift barrier”



“Fragmentation barrier”



...BUT!



Need dust traps:

"Radial drift barrier"

- dead zones

- **snow lines**

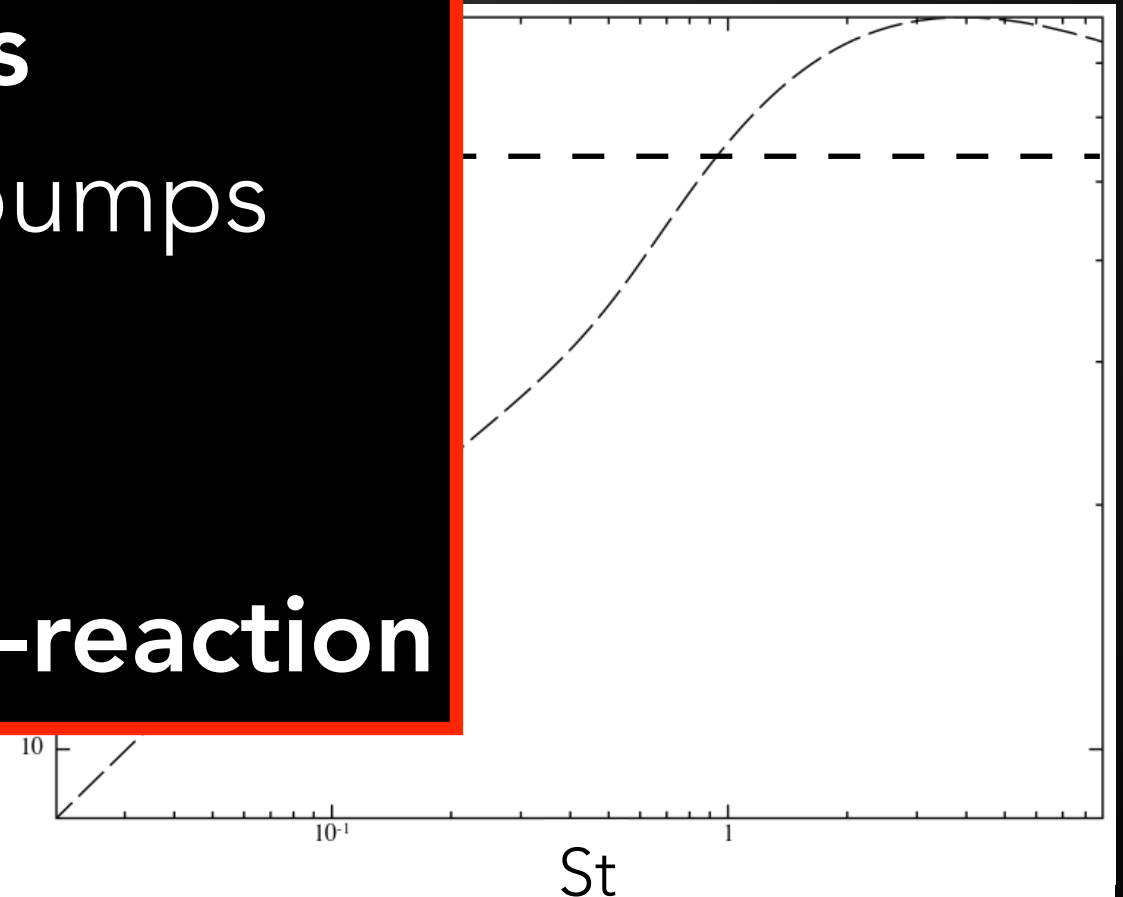
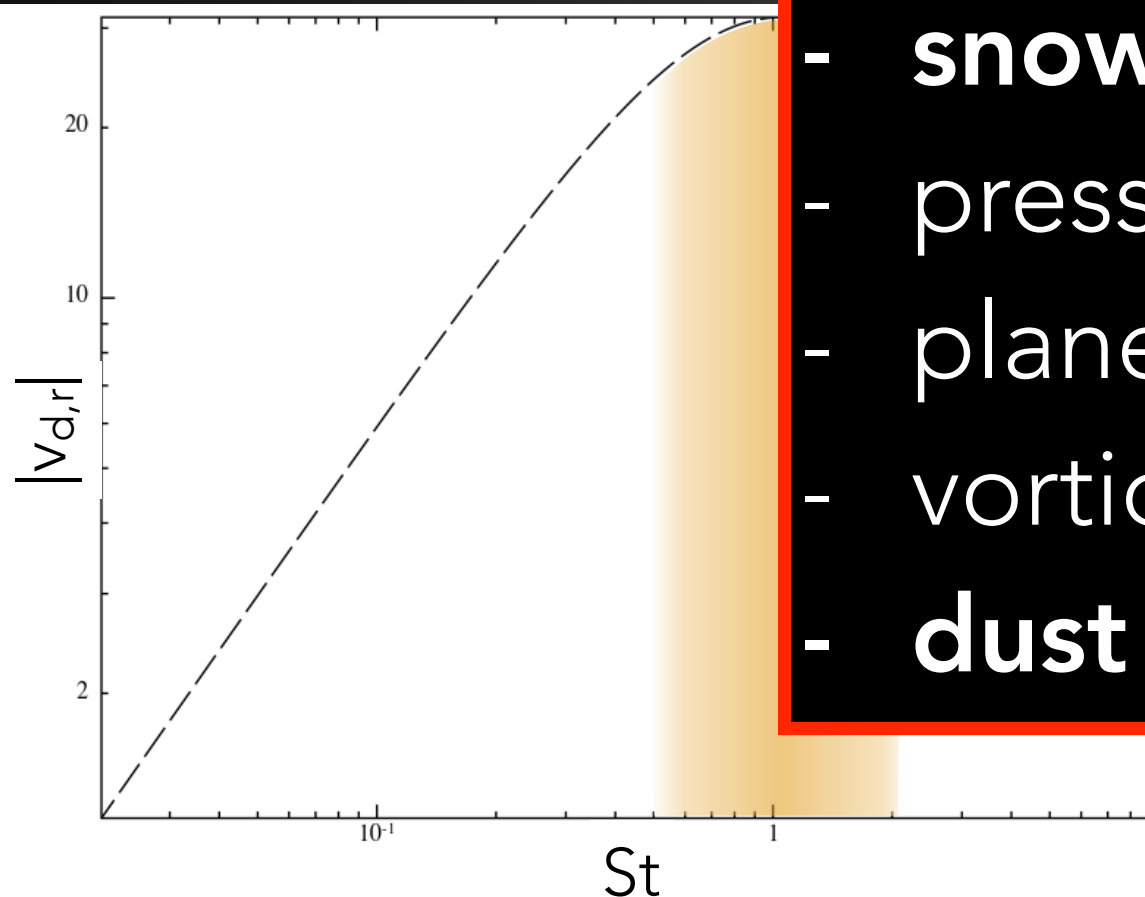
- pressure bumps

- planets

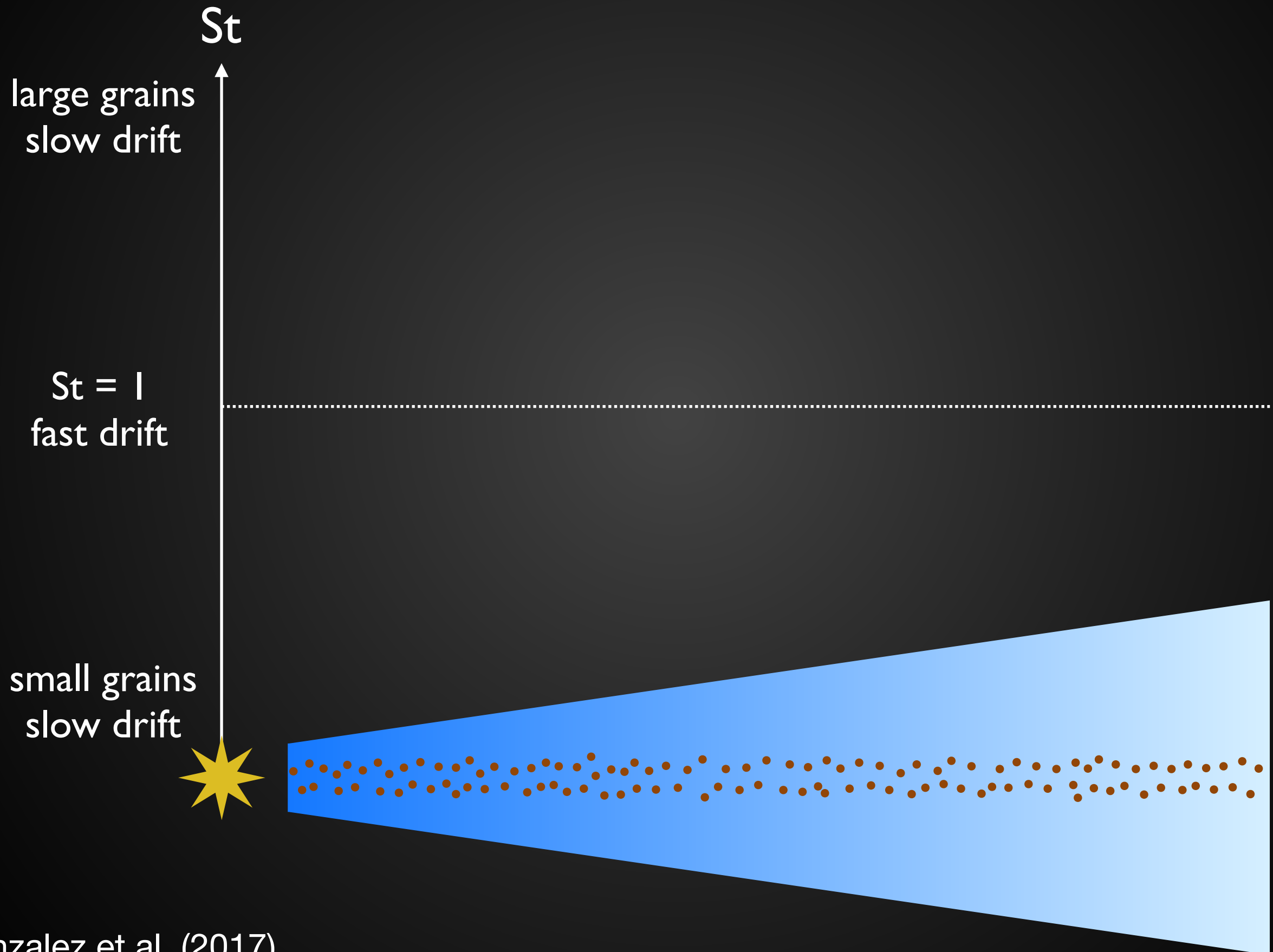
- vortices

- **dust back-reaction**

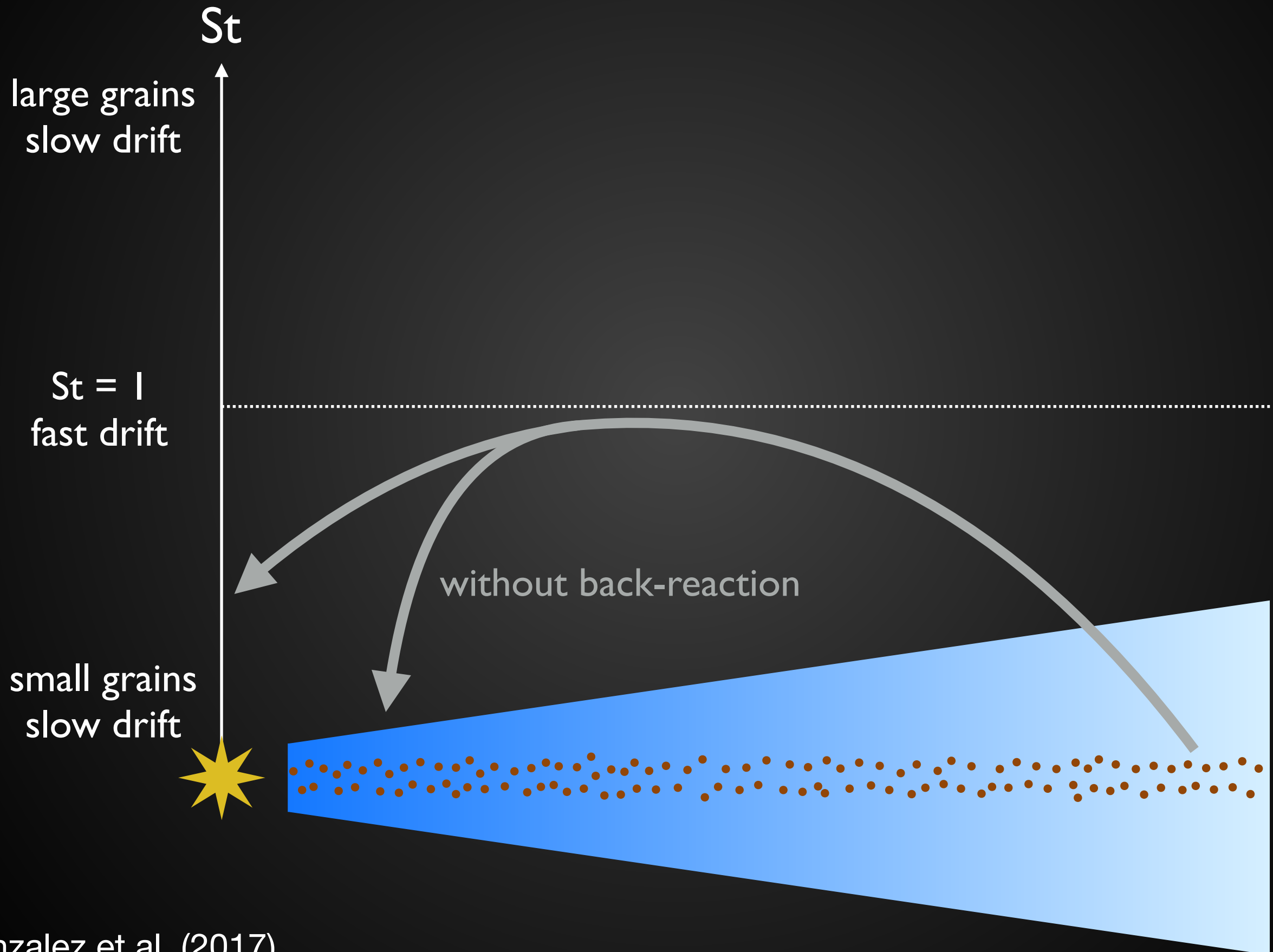
"Accretion barrier"



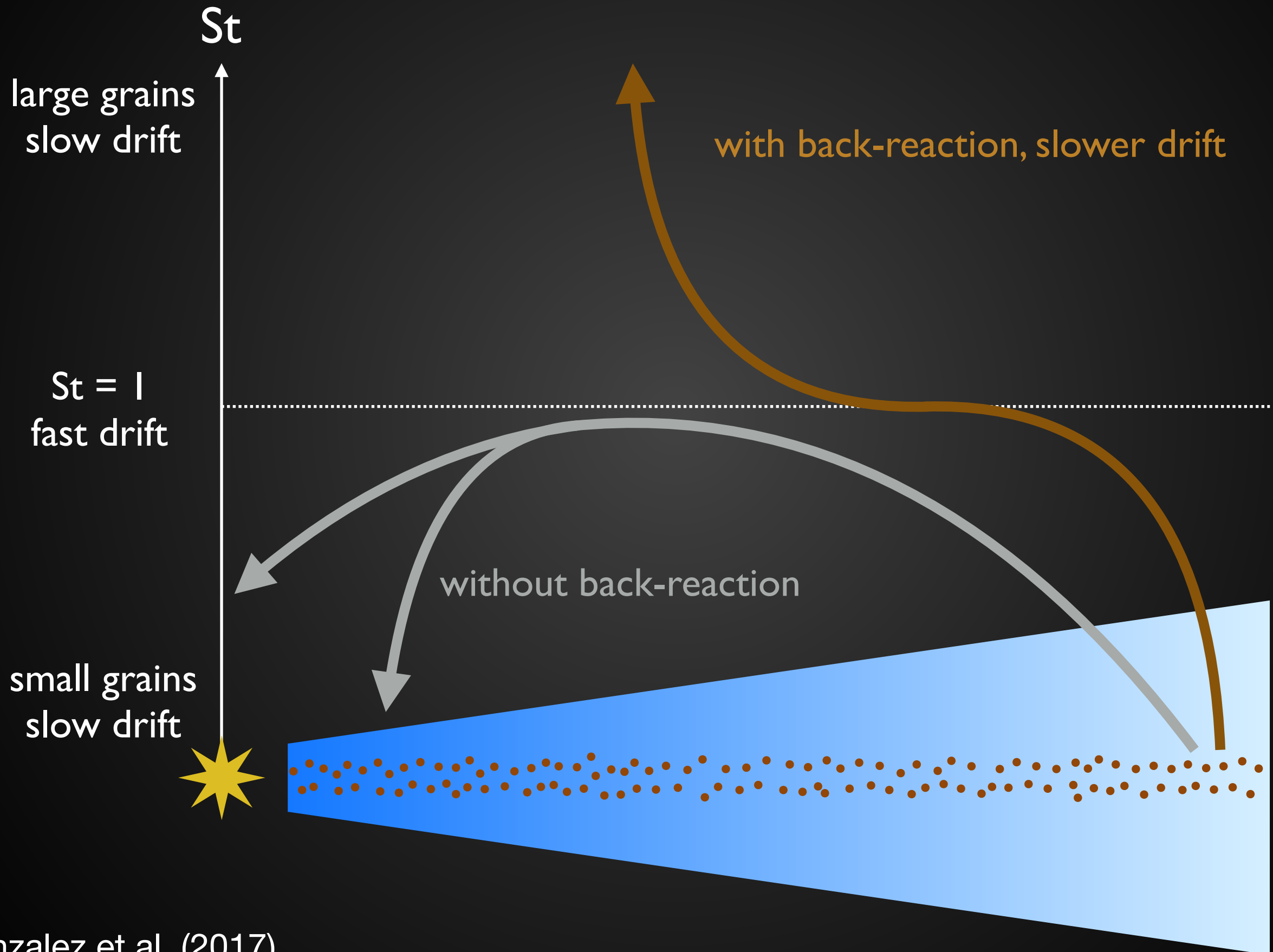
SELF-INDUCED DUST TRAP



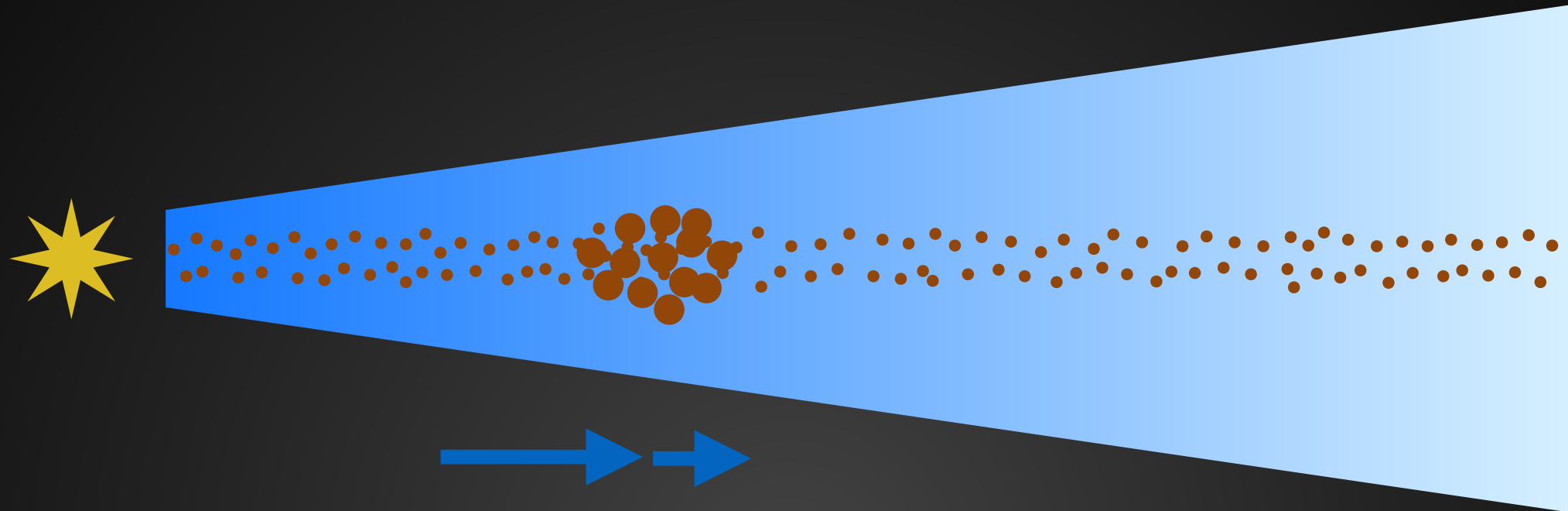
SELF-INDUCED DUST TRAP



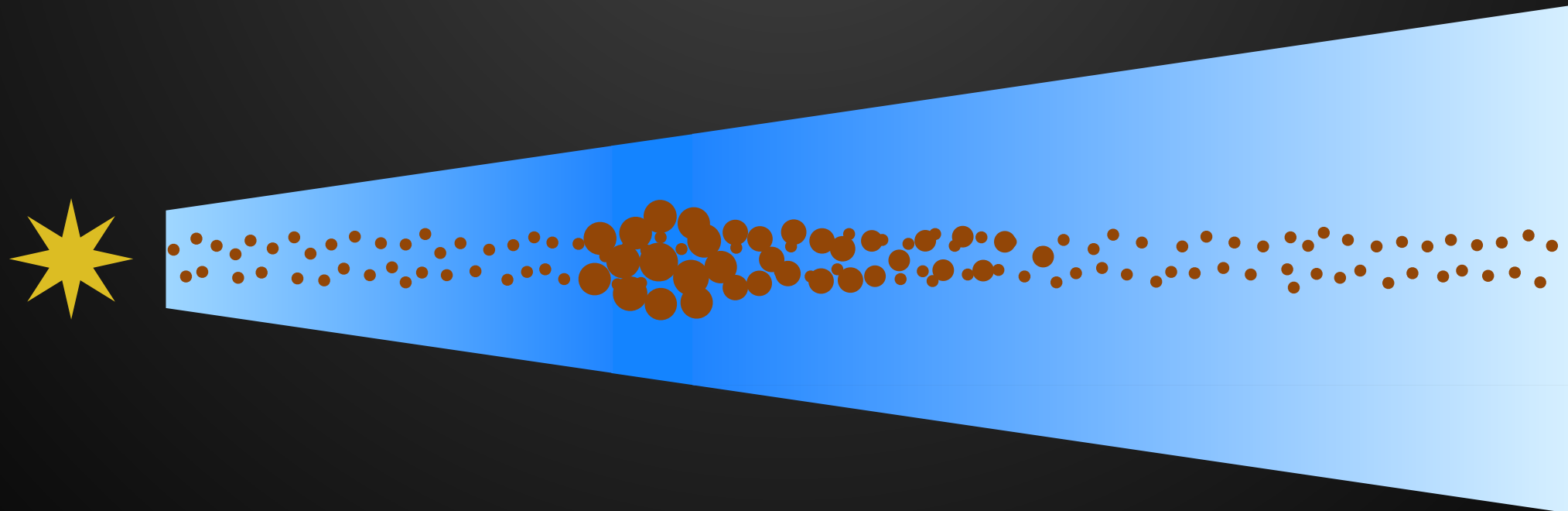
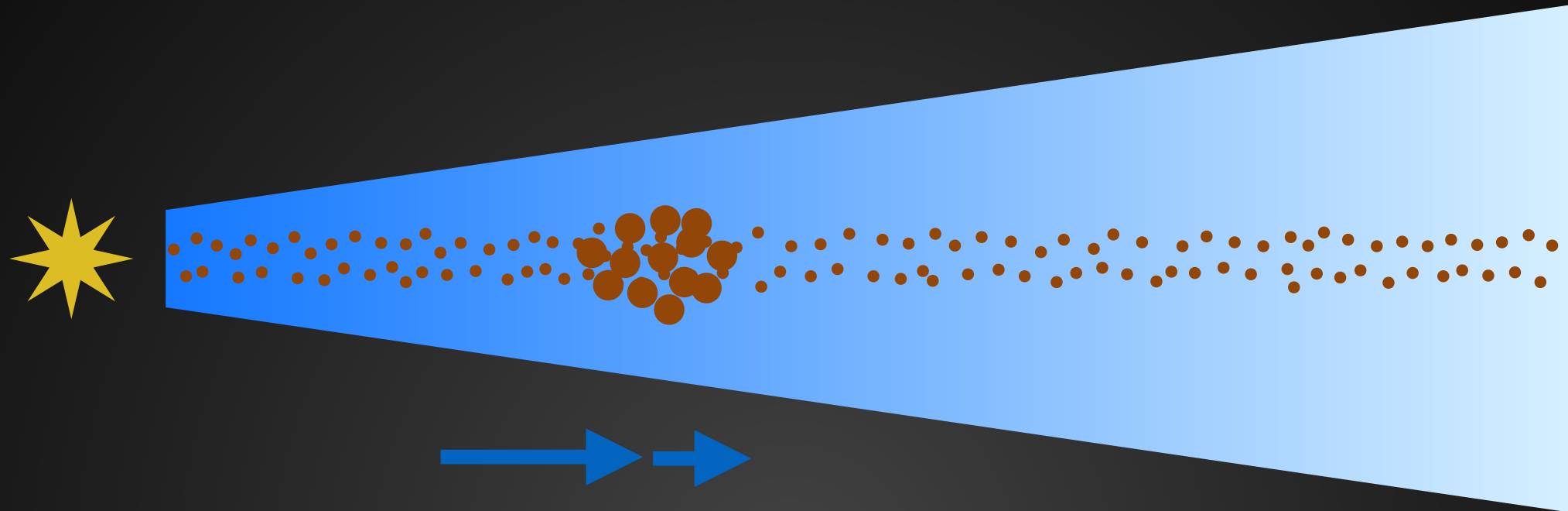
SELF-INDUCED DUST TRAP



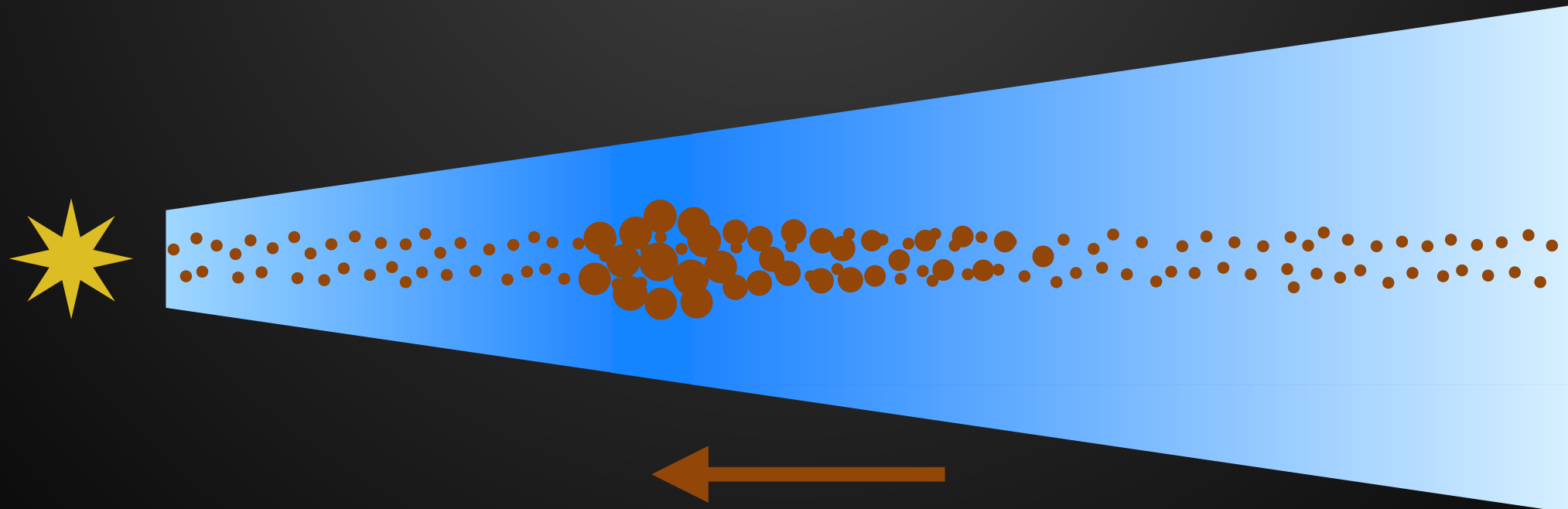
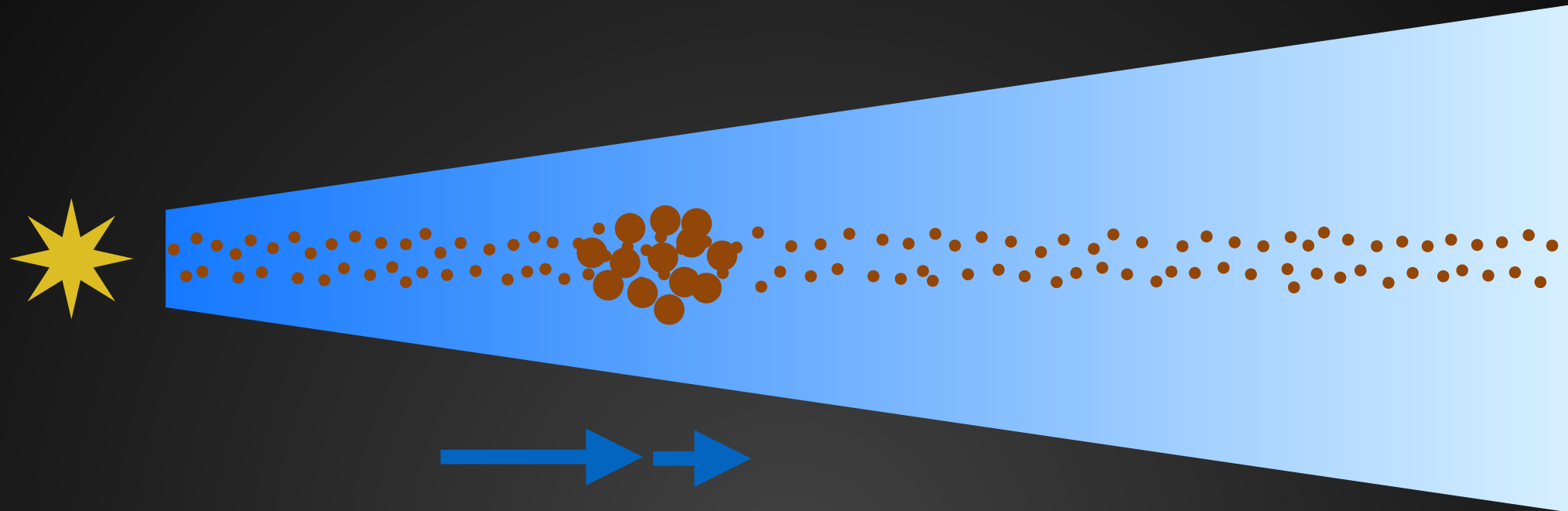
SELF-INDUCED DUST TRAP



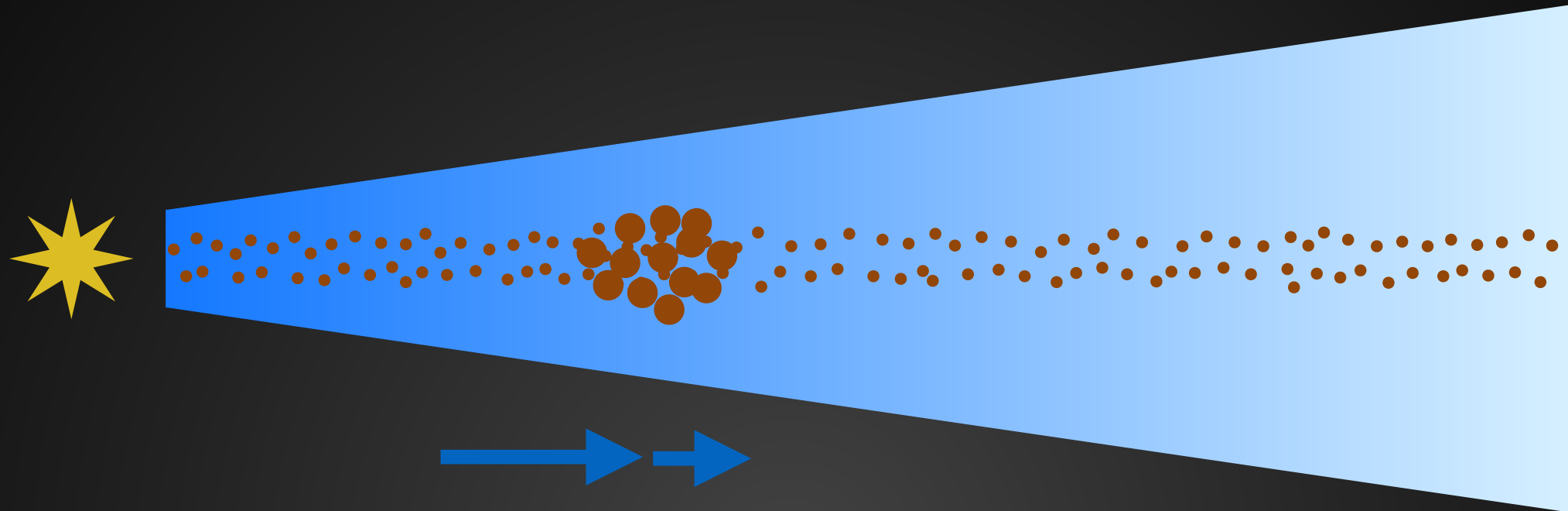
SELF-INDUCED DUST TRAP



SELF-INDUCED DUST TRAP



SELF-INDUCED DUST TRAP



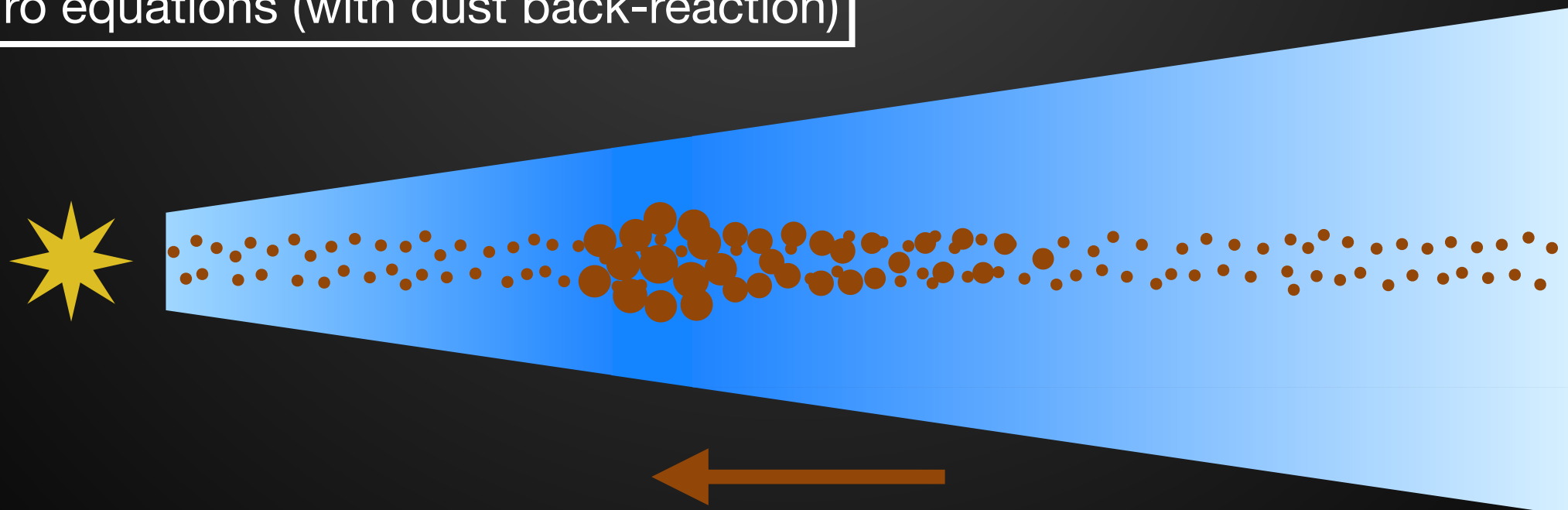
Lyon SPH code:

3D gas+dust (two-fluid)
growth and fragmentation
solve hydro equations (with dust back-reaction)

Barrière-fouchet et al. (2007)

Laibe et al. (2008)

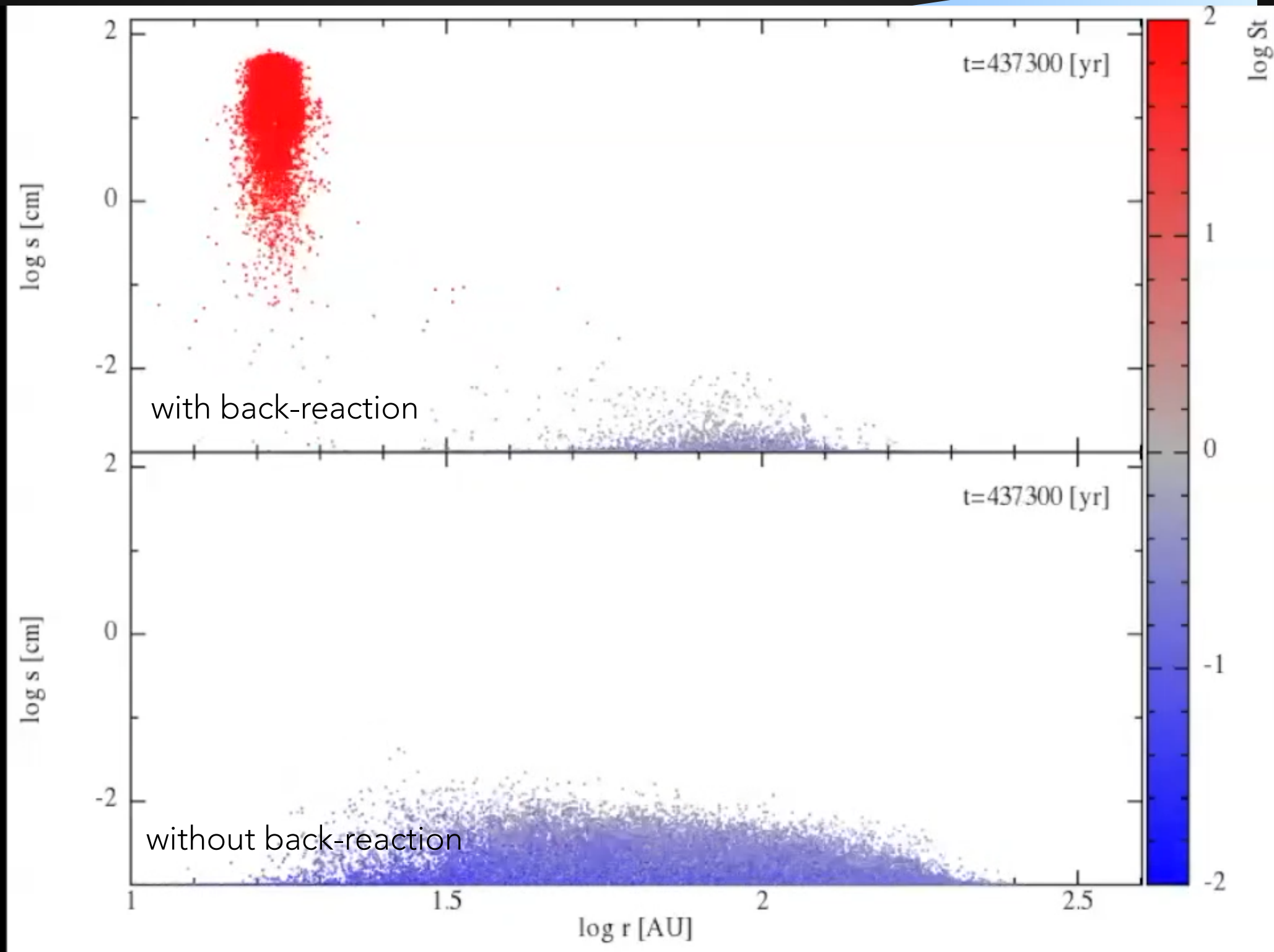
Gonzalez et al. (2015)



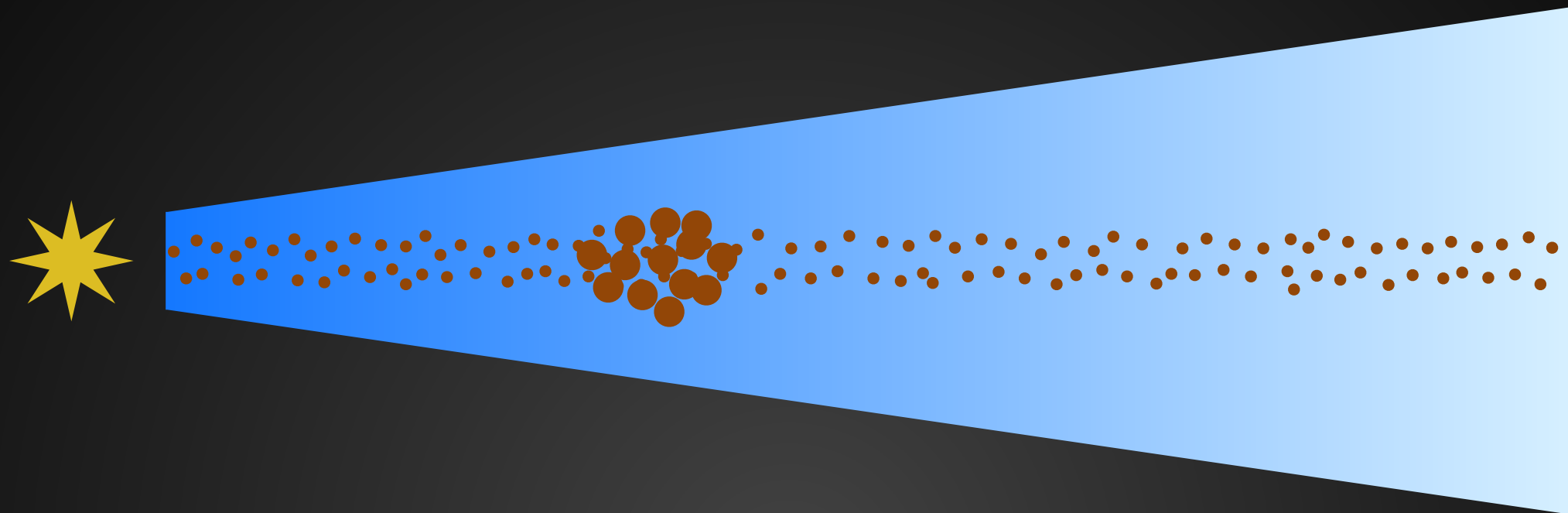
Gonzalez et al. (2017)

SELF-INDUCED DUST TRAP

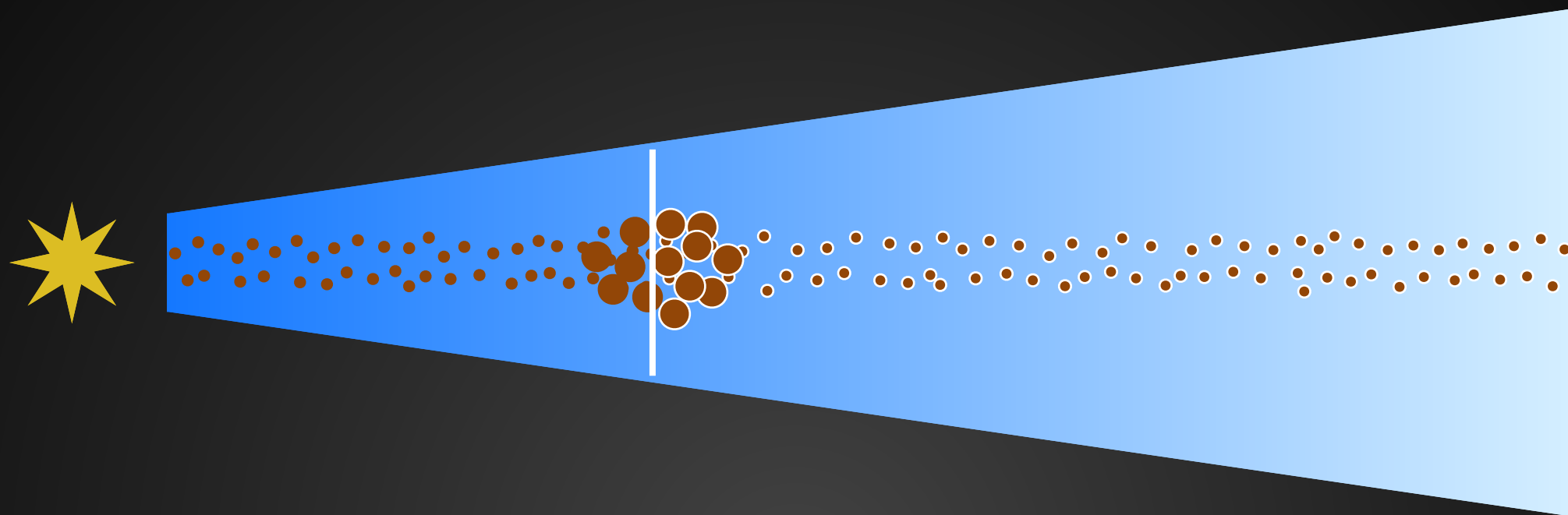
$$V_{\text{frag}} = 15 \text{ m} \cdot \text{s}^{-1}$$



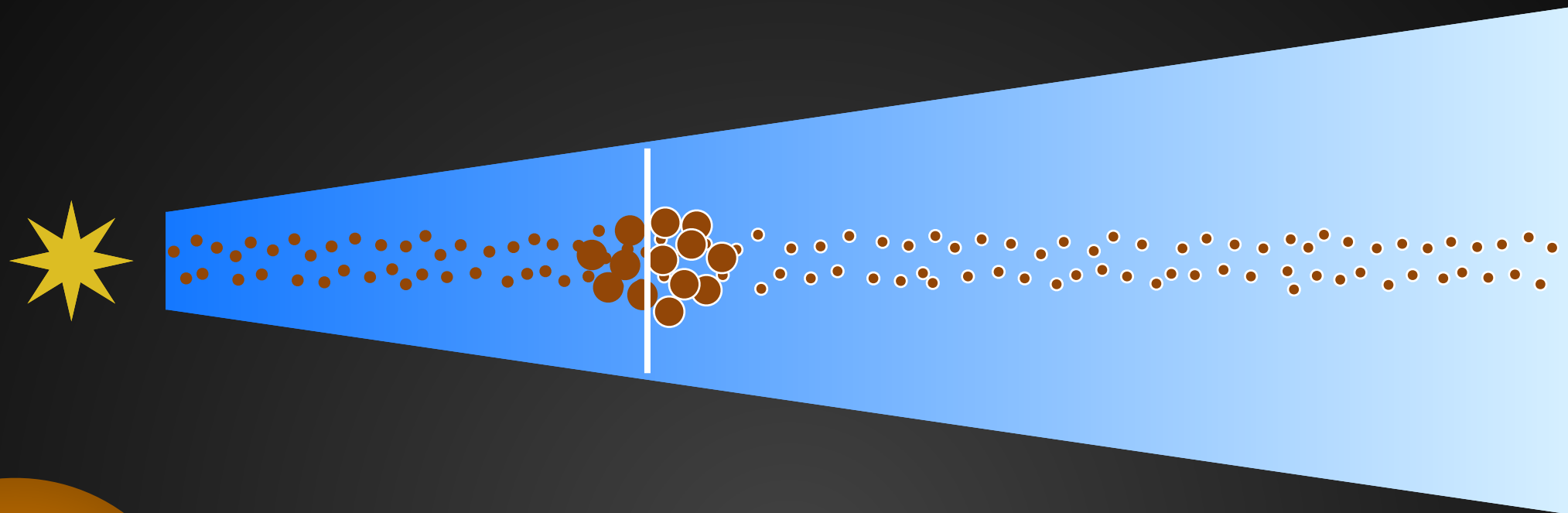
SNOW LINES ?



SNOW LINES ?



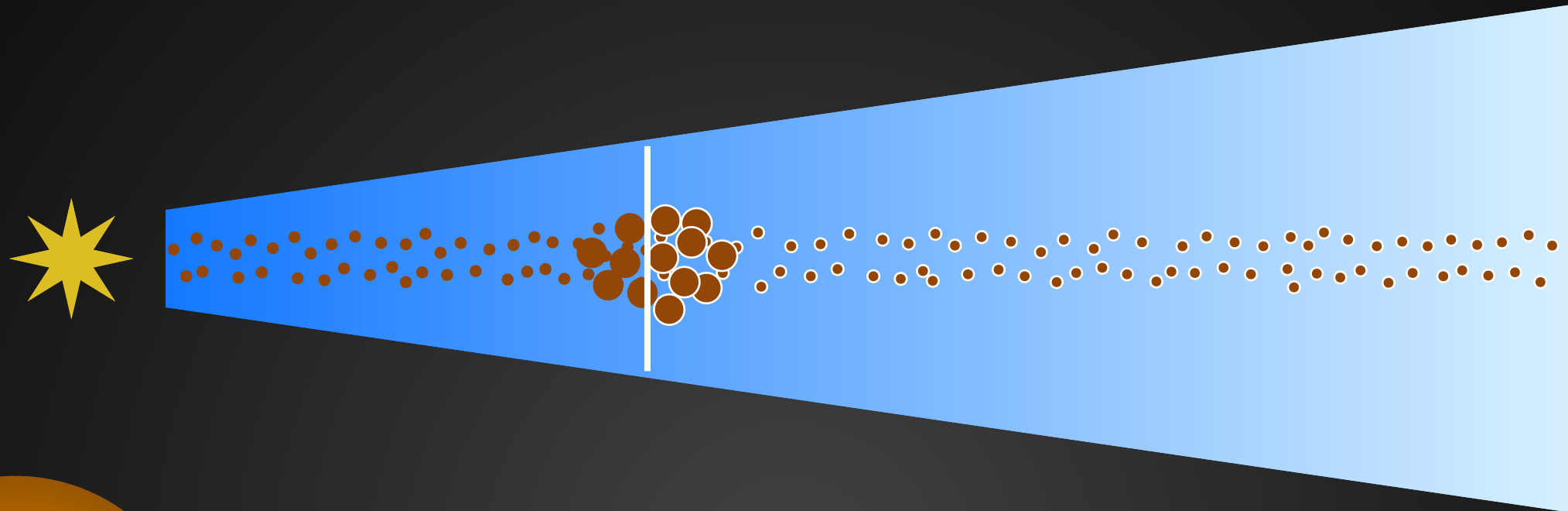
SNOW LINES ?



BARE
SILICATE
CORE

sticks poorly: $V_{\text{frag}} \sim 1 \text{ m.s}^{-1}$

SNOW LINES ?



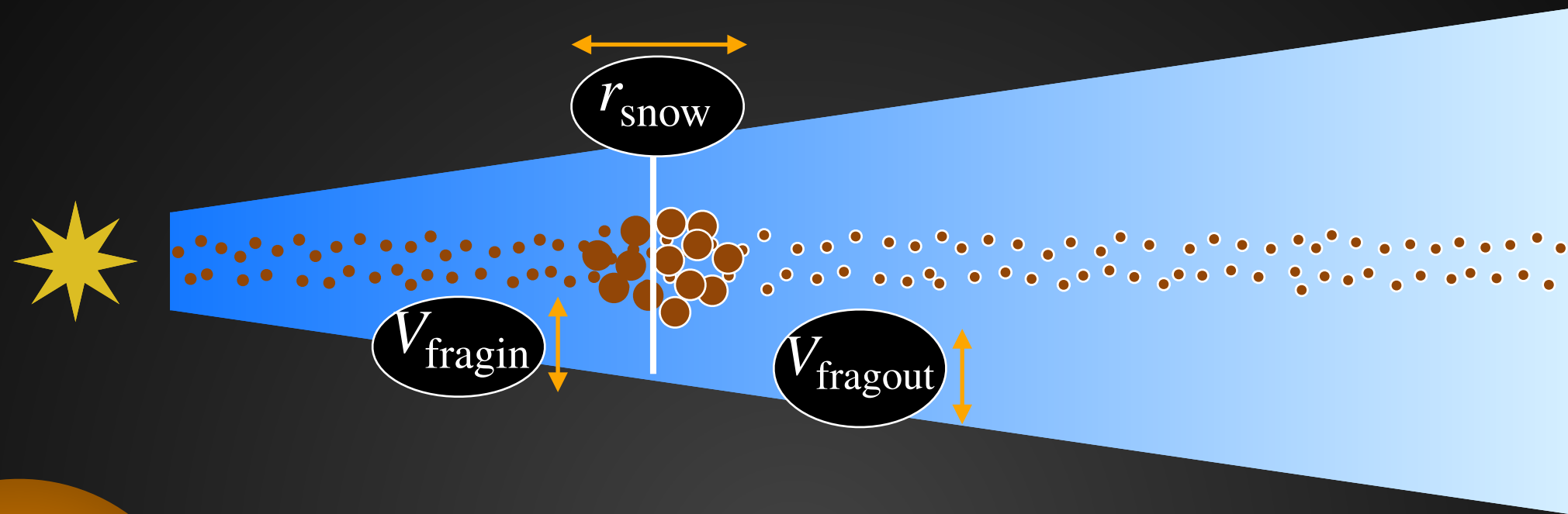
BARE
SILICATE
CORE

sticks poorly: $V_{\text{frag}} \sim 1 \text{ m.s}^{-1}$

SILICATE
CORE
+
H₂O MANTLE

sticks well: $V_{\text{frag}} \sim 15 - 20 \text{ m.s}^{-1}$

SNOW LINES ?



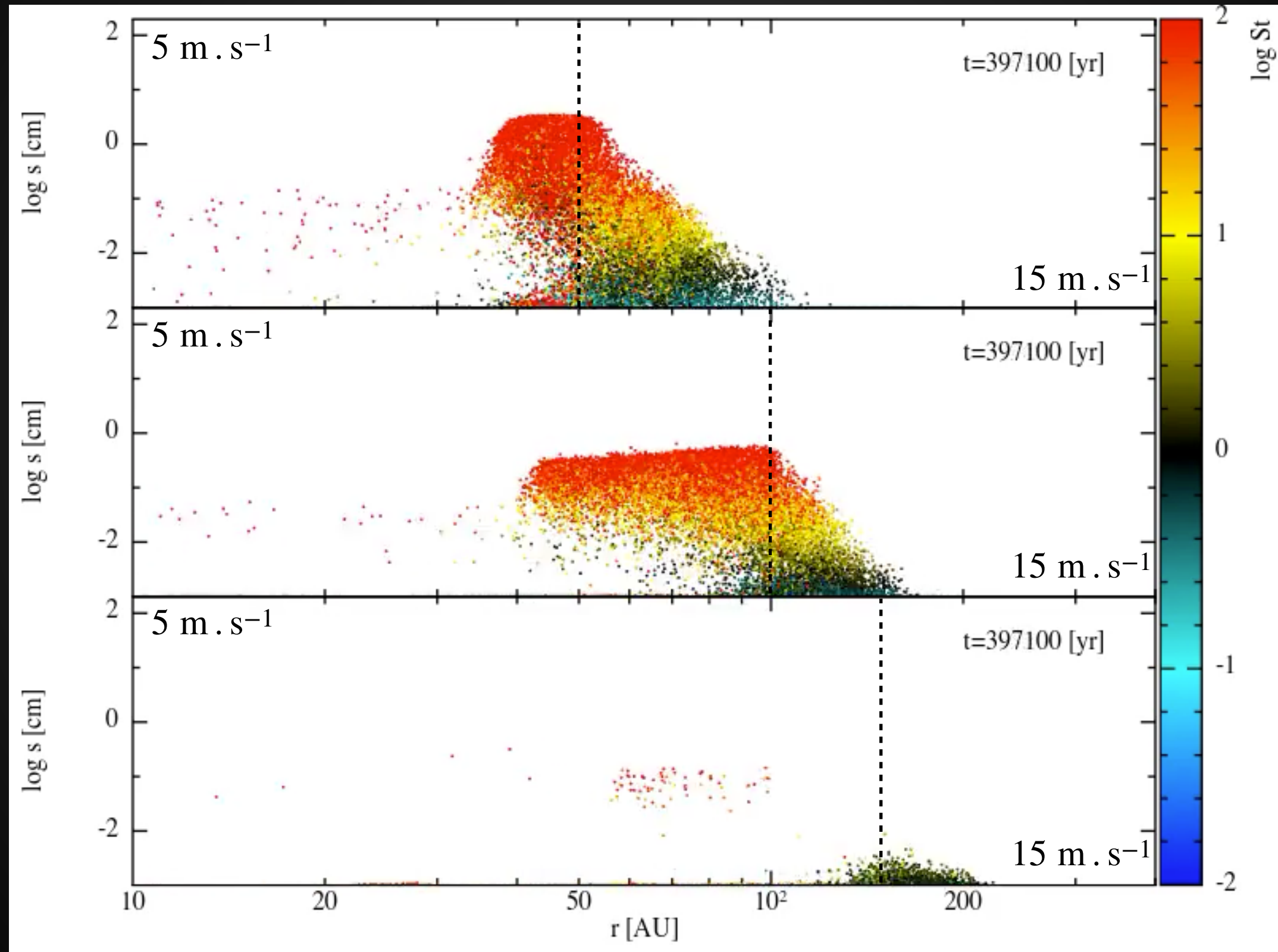
BARE
SILICATE
CORE

sticks poorly: $V_{\text{frag}} \sim 1 \text{ m.s}^{-1}$

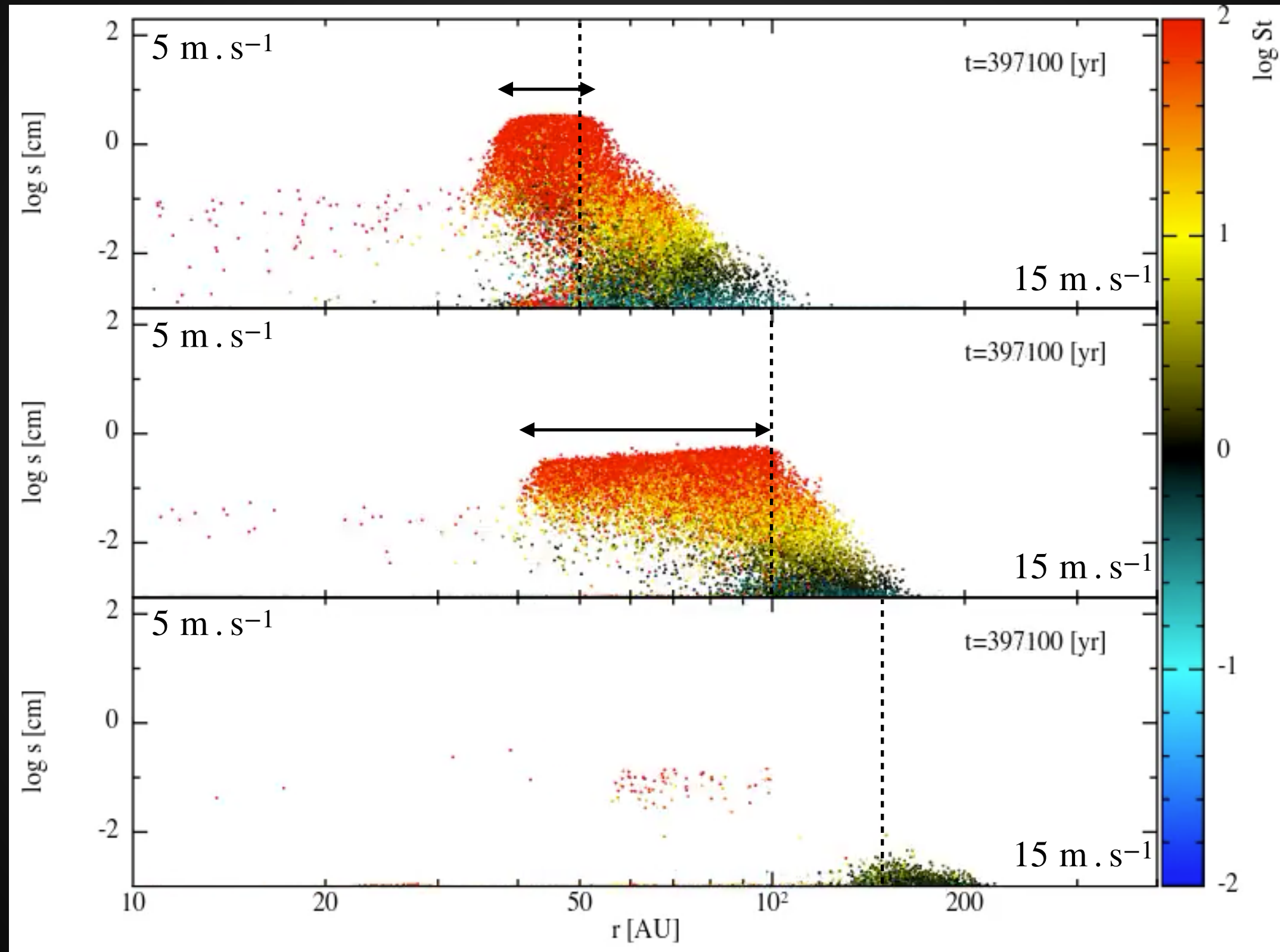
SILICATE
CORE
+
 H_2O MANTLE

sticks well: $V_{\text{frag}} \sim 15 - 20 \text{ m.s}^{-1}$

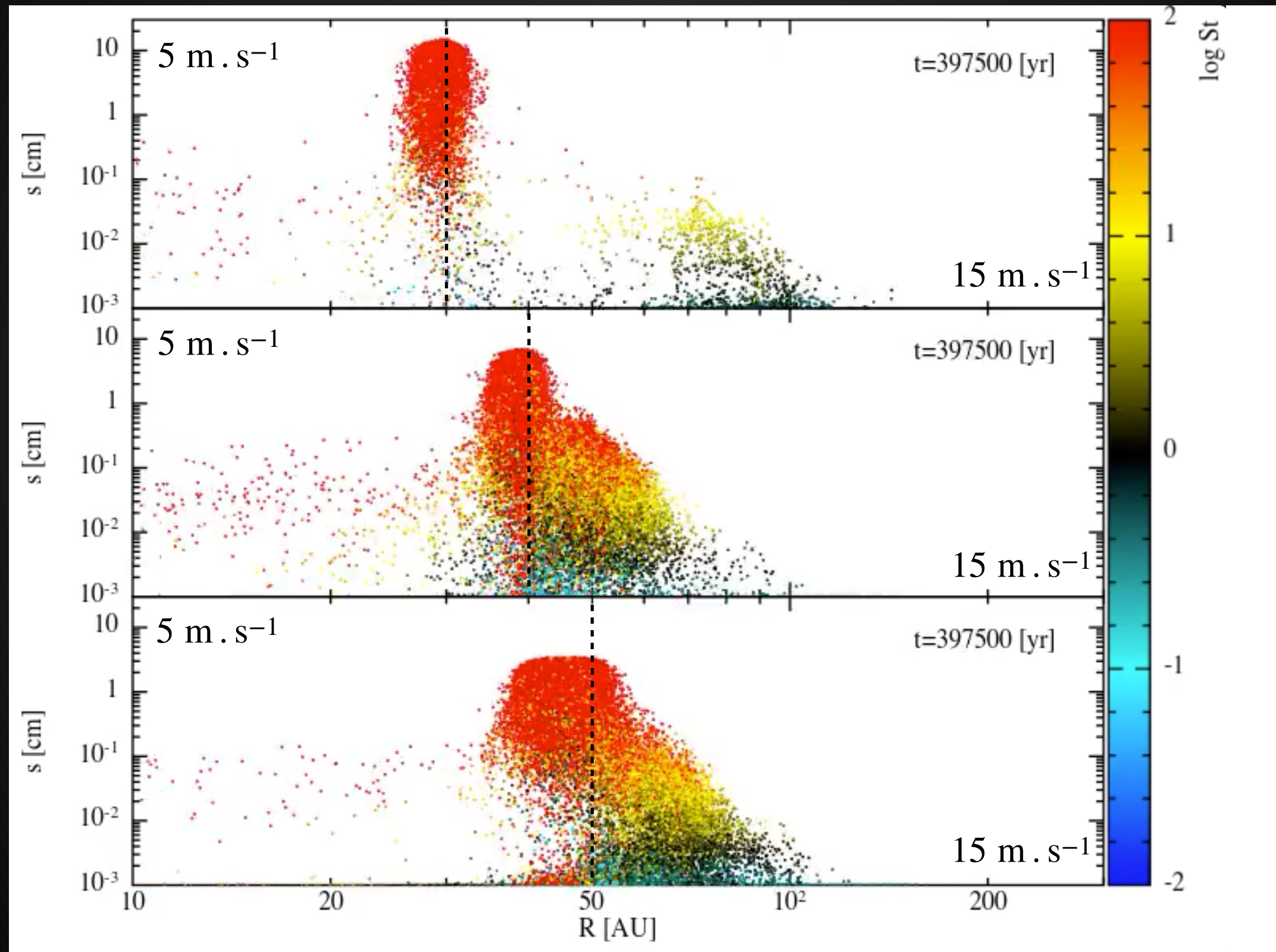
SNOW LINE POSITION



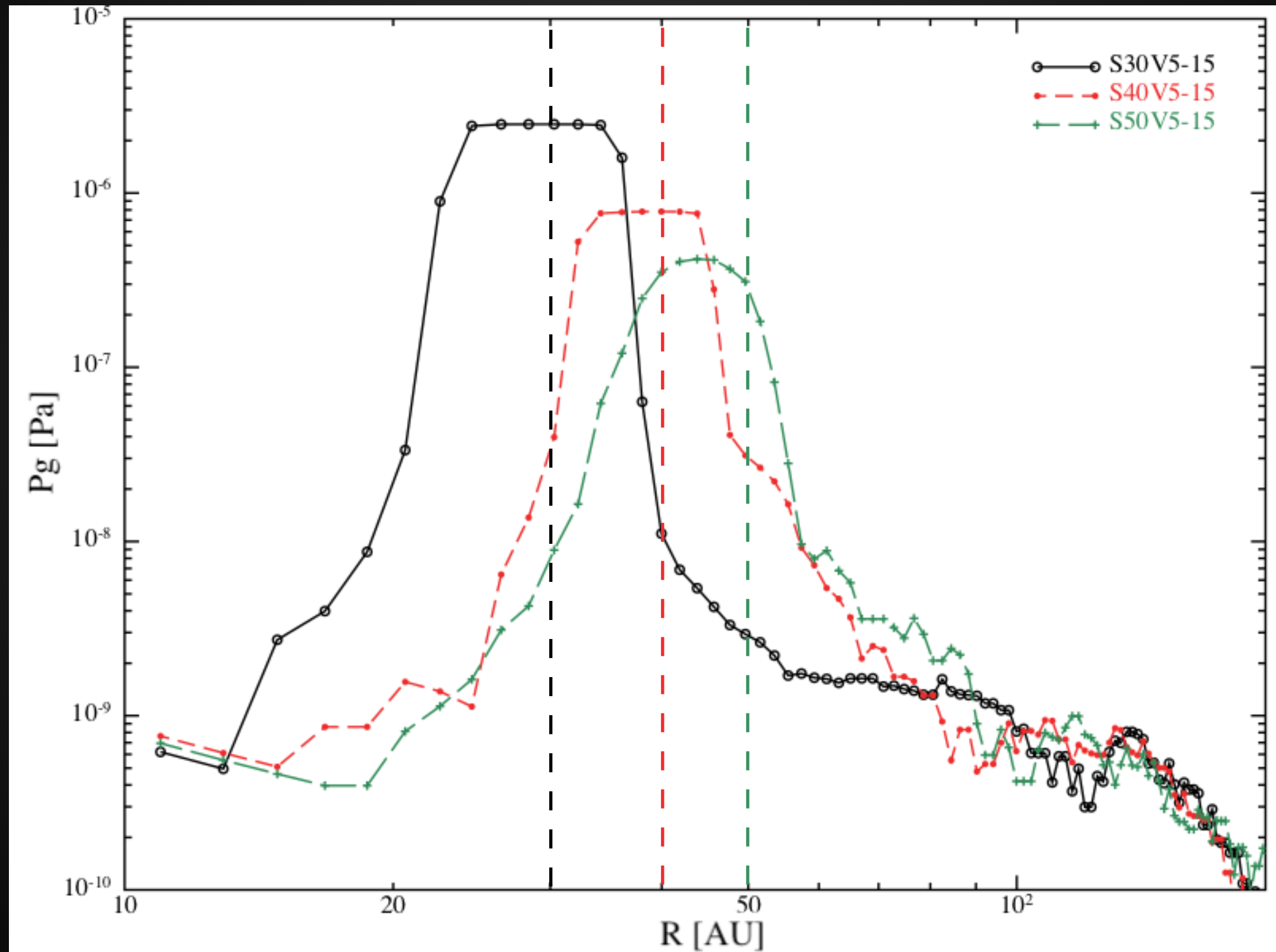
SNOW LINE POSITION



SNOW LINE POSITION



SNOW LINE POSITION

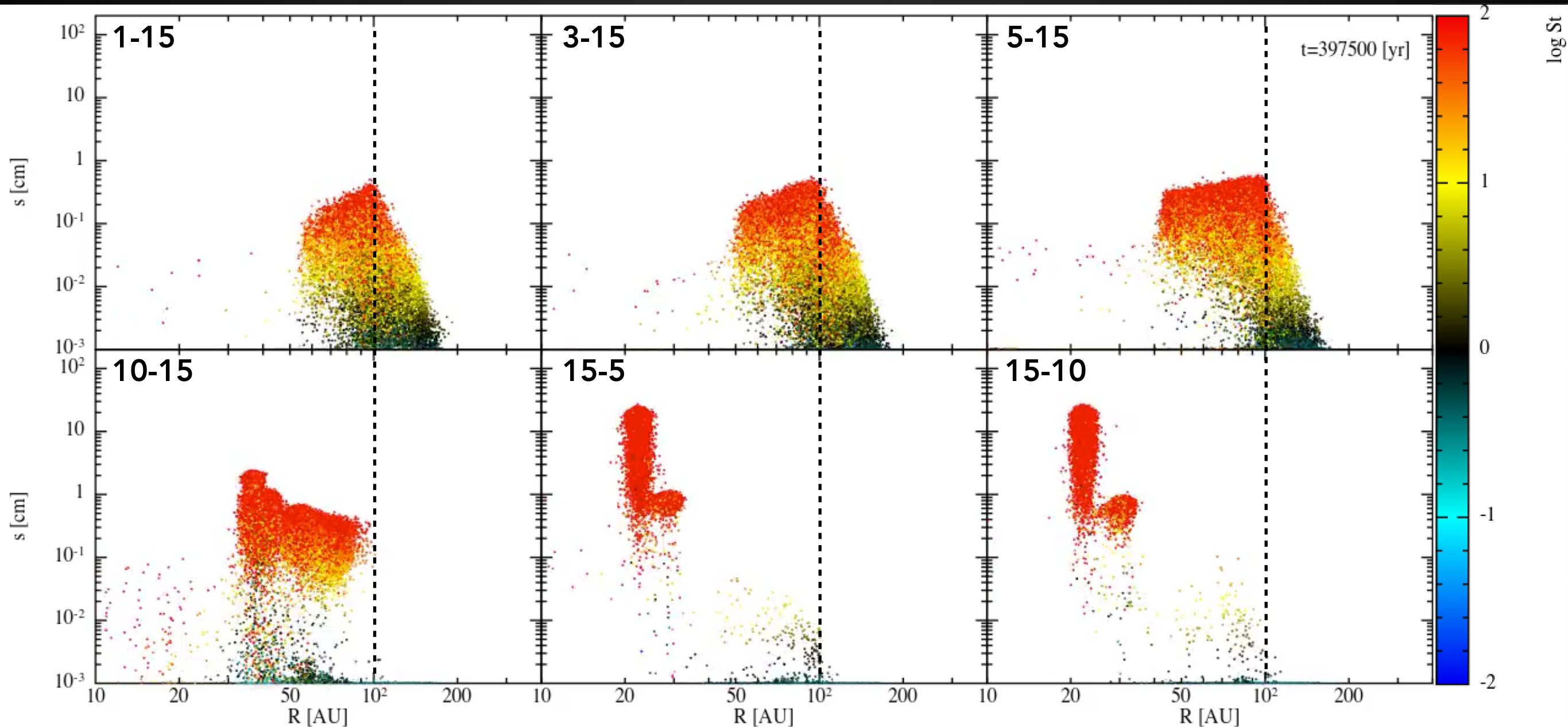


CO SNOW LINE

$$T \propto r^{-q} \rightarrow r_{\text{snow}} \propto T_{\text{subli}}^{-1/q}$$

$$T_{\text{subli}}(\text{CO}) \sim 20 \text{ K}$$

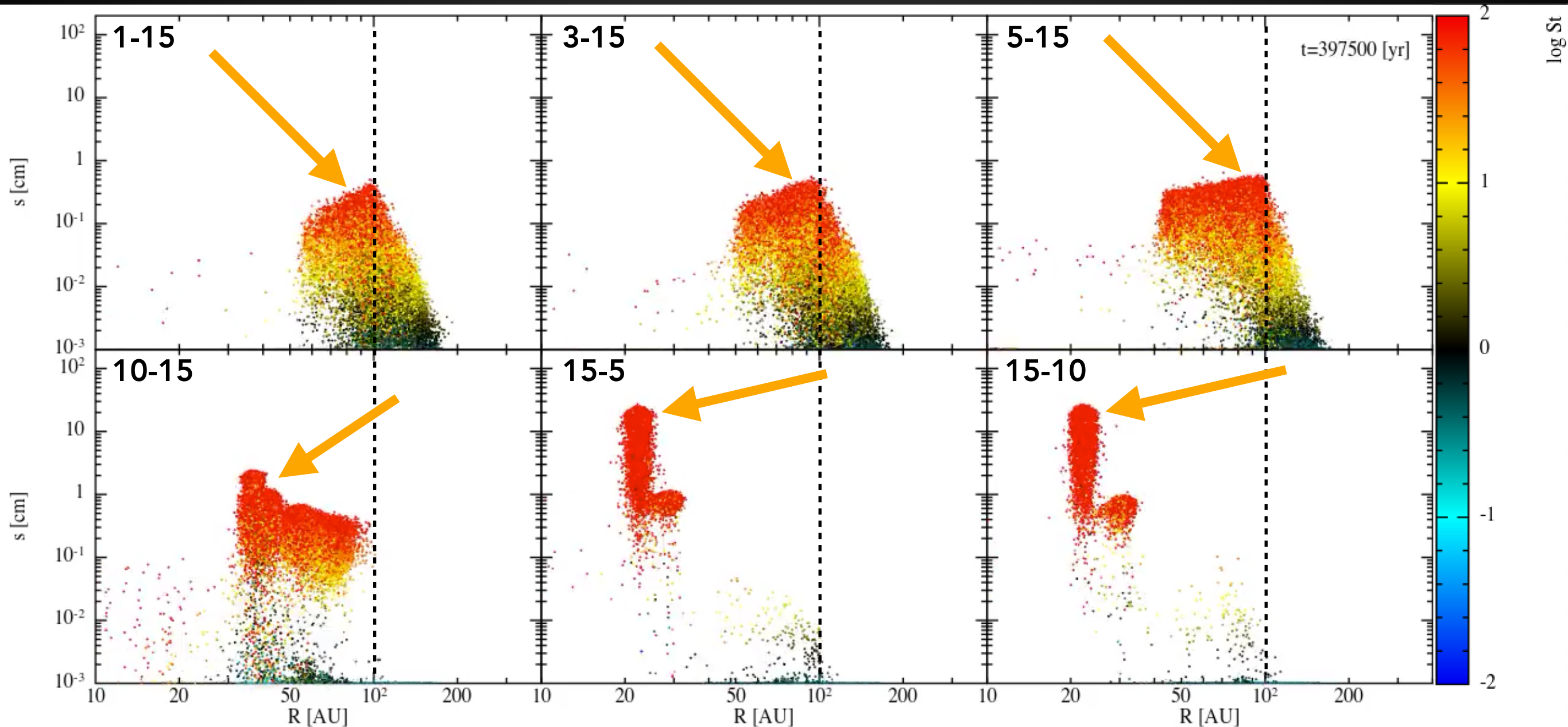
CO SNOW LINE



$$T_{\text{subli}}(\text{CO}) \sim 20 \text{ K}$$

$$T \propto r^{-q} \rightarrow r_{\text{snow}} \propto T_{\text{subli}}^{-1/q}$$

CO SNOW LINE

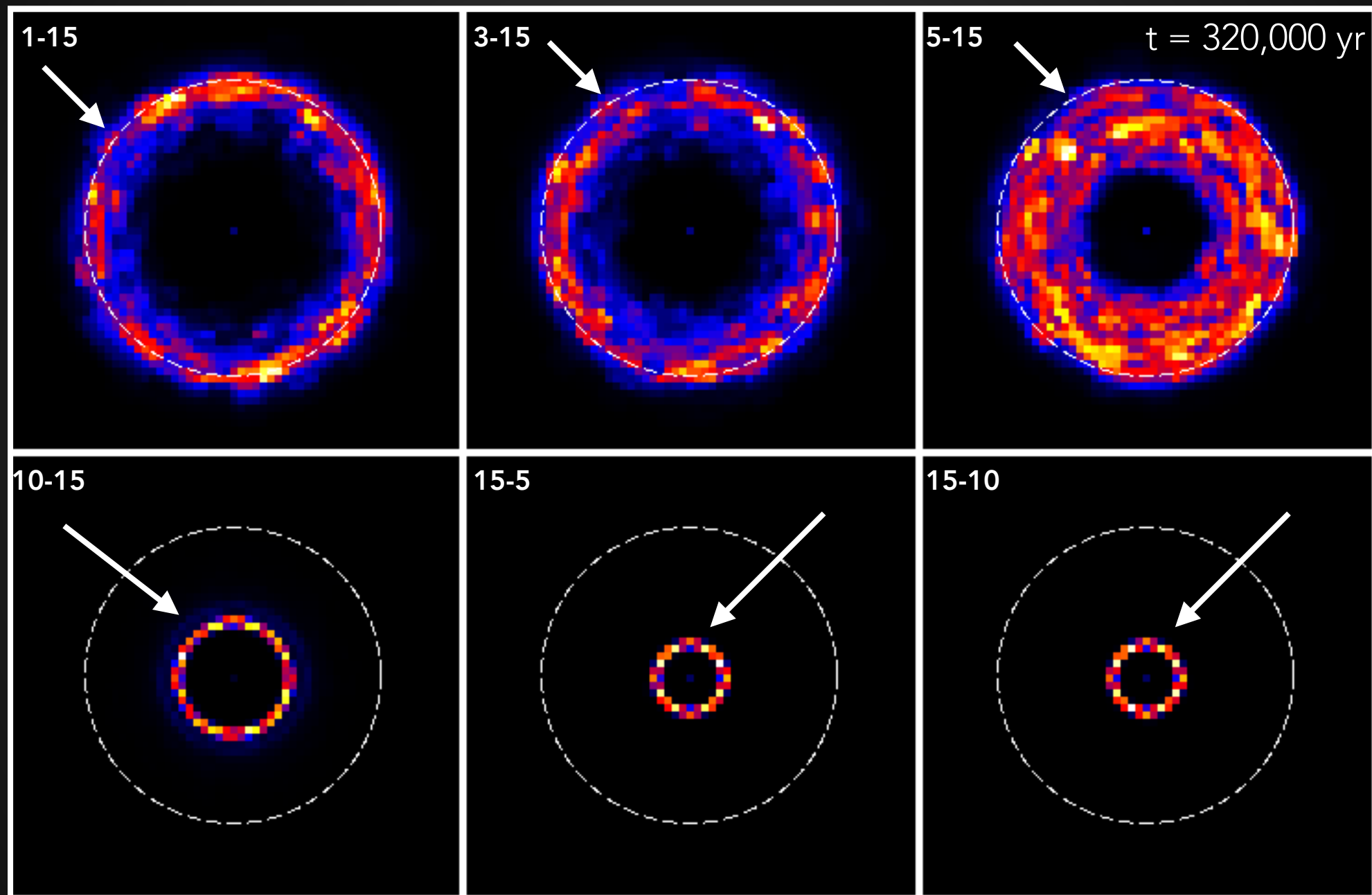


$T_{\text{subli}}(\text{CO}) \sim 20$ K

$$T \propto r^{-q} \rightarrow r_{\text{snow}} \propto T_{\text{subli}}^{-1/q}$$

Vericel & Gonzalez (in prep)

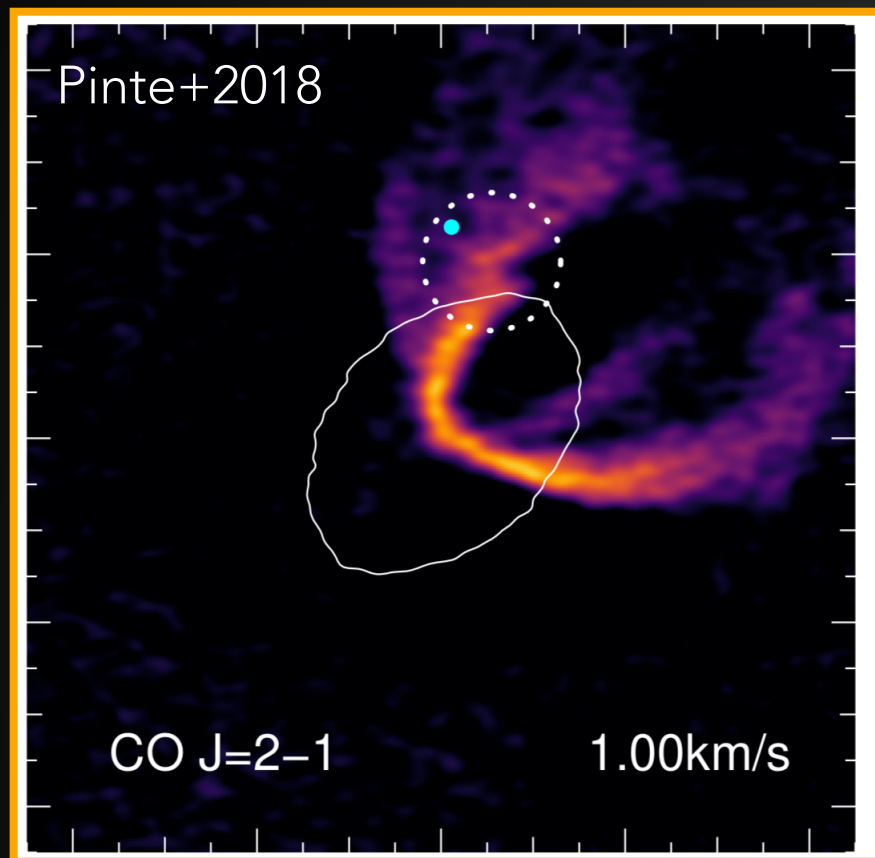
OBSERVATIONAL SIGNATURES



$850 \mu\text{m}, i=0^\circ$

TAKE AWAY MESSAGES

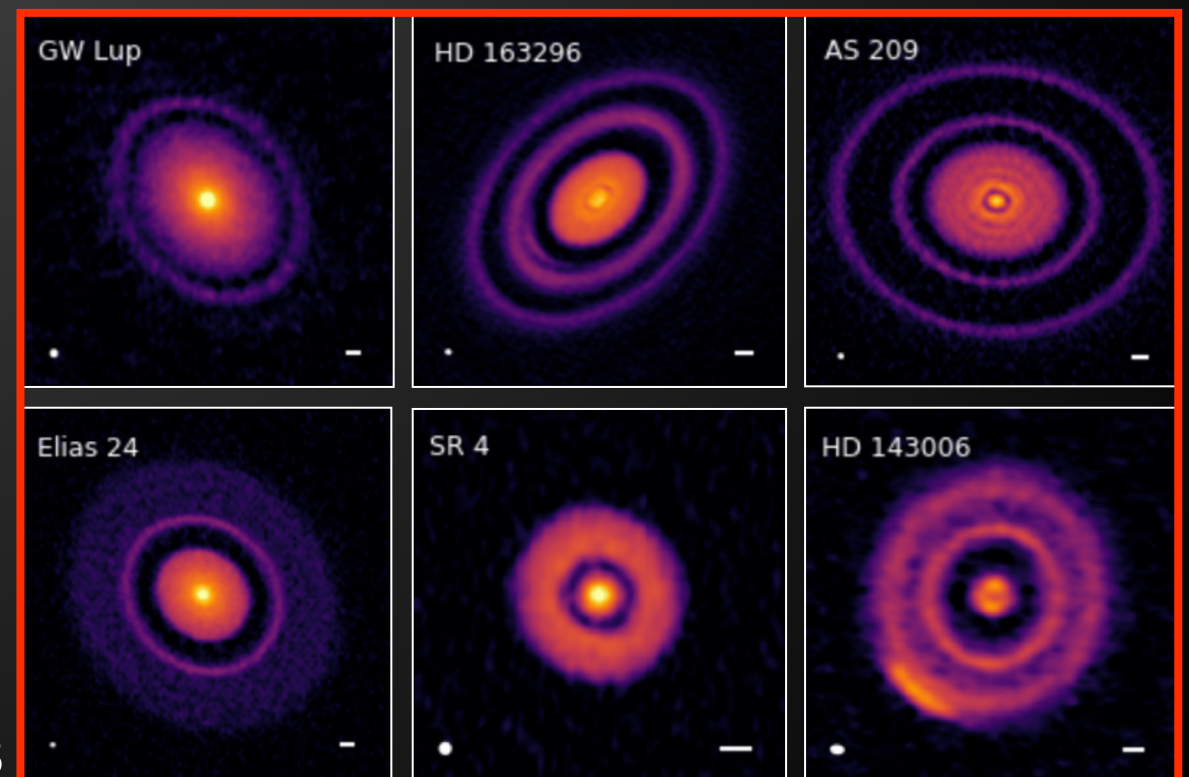
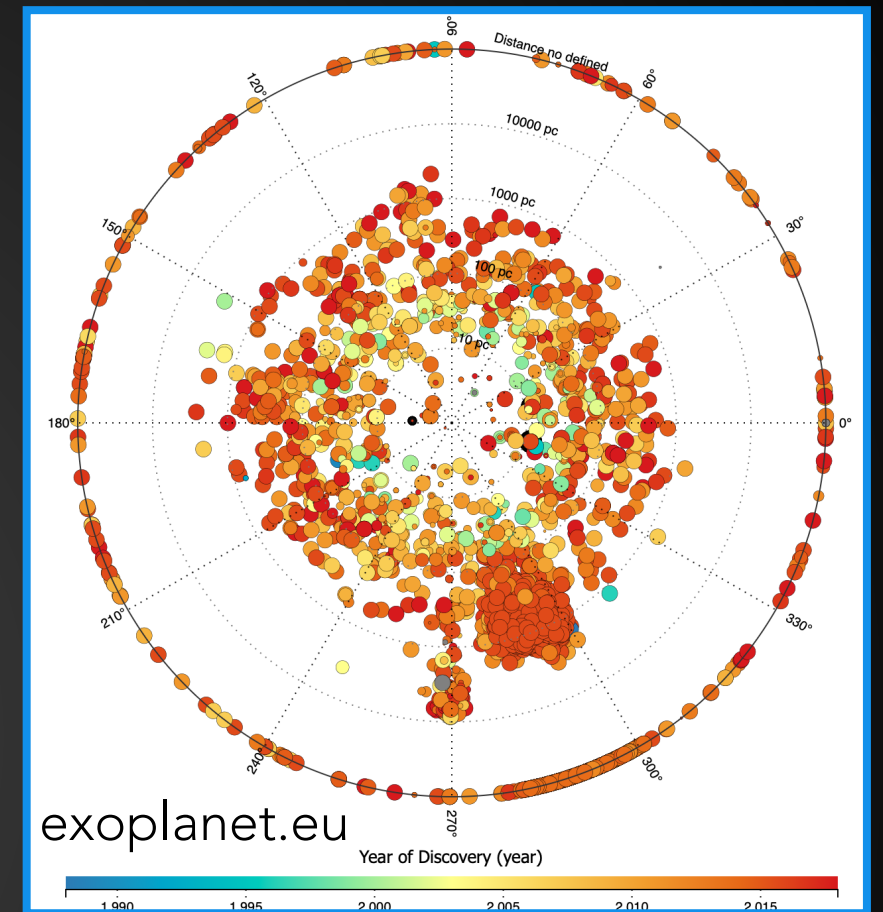
Self-induced
dust trap formation
is a robust mechanism,
even with snowlines.



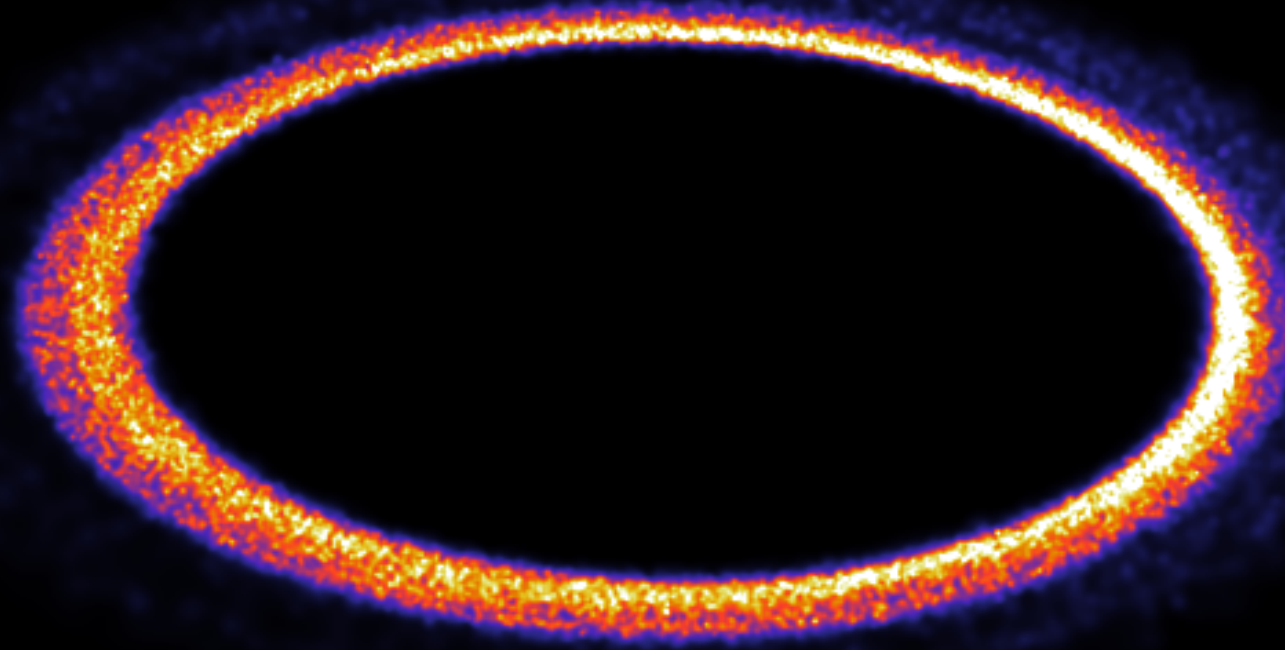
The timescale of these traps formation
is of a few hundred thousand years
<< disc lifetime.

It could explain some of
the rings/structures seen with ALMA.
(in progress)

Andrews+2018



THANK YOU



SF2A 2019, Nice



Arnaud Vericel, Jean-François Gonzalez
Centre de Recherche Astrophysique de Lyon